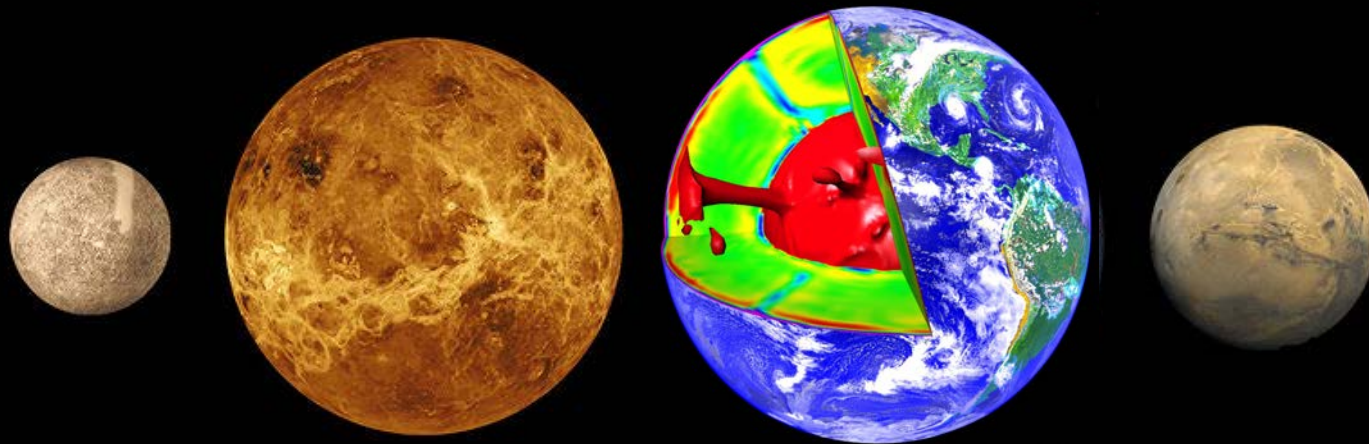


# Self-consistent plate tectonics in global convection models: Recent progress

Paul J. Tackley,

Diogo Lourenco, Antoine Rozel, ETH Zürich

Claire Mallard & Nicolas Coltice, ENS Lyon



# Overview

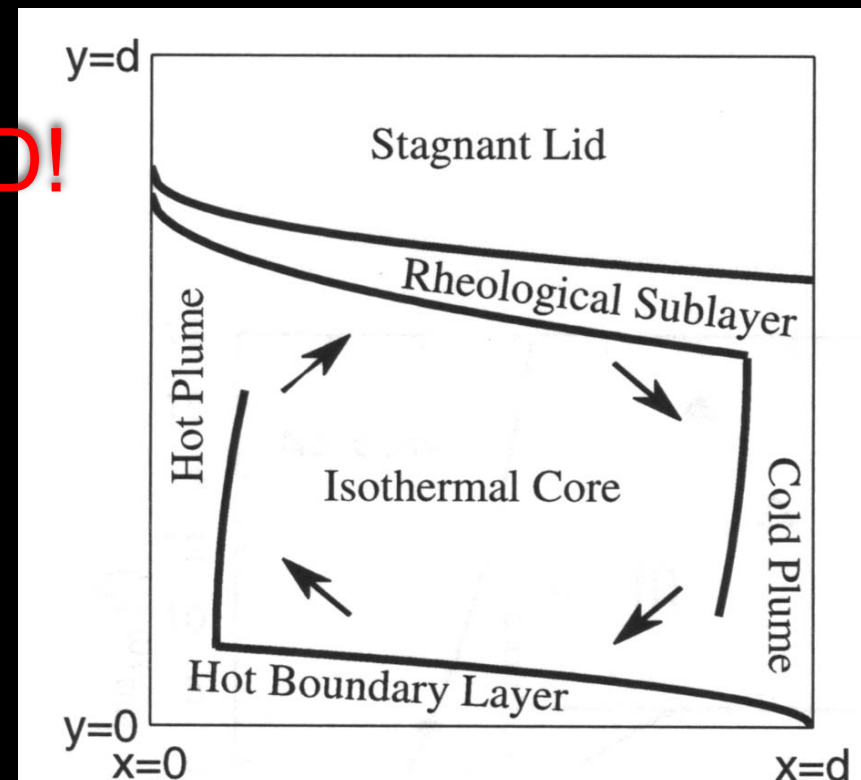
- Review of the problem
- Global statistics: easy to get correct?
- Magmatism->crust is important!

**Why does Earth have  
plate tectonics?**

# The plate problem

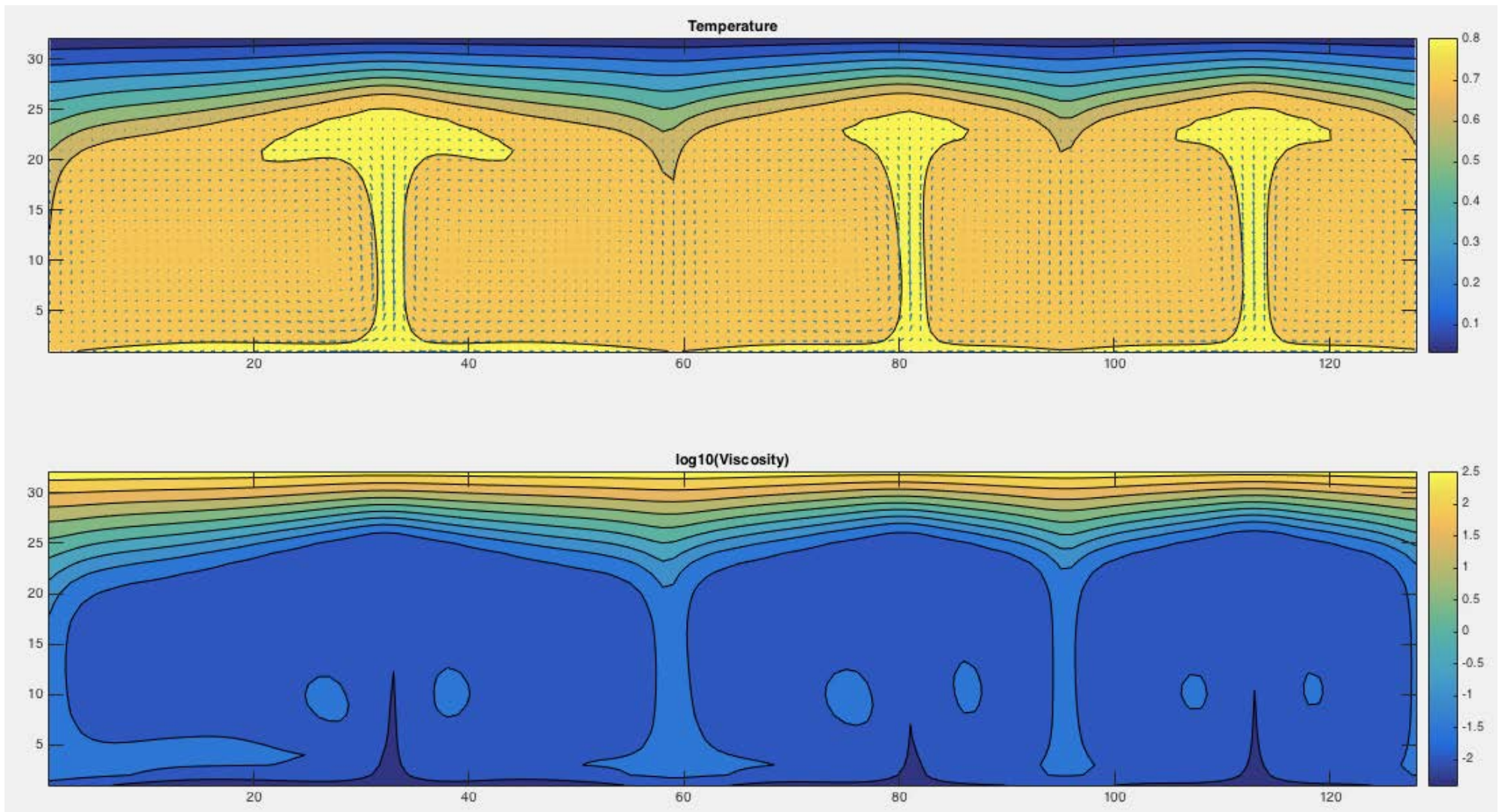
- Viscous, T-dependent rheology appropriate for the mantle leads to a stagnant lid
- $\exp(E/kT)$  where  $E \sim 340$  kJ/mol
- T from 1600  $\rightarrow$  300 K
- $\Rightarrow 1.3 \times 10^{48}$  variation
- $\Rightarrow$  **RIGID/STAGNANT LID!**

Only small  $\Delta T$  participates in convection: enough to give  $\Delta\eta$  factor  $\sim 10$

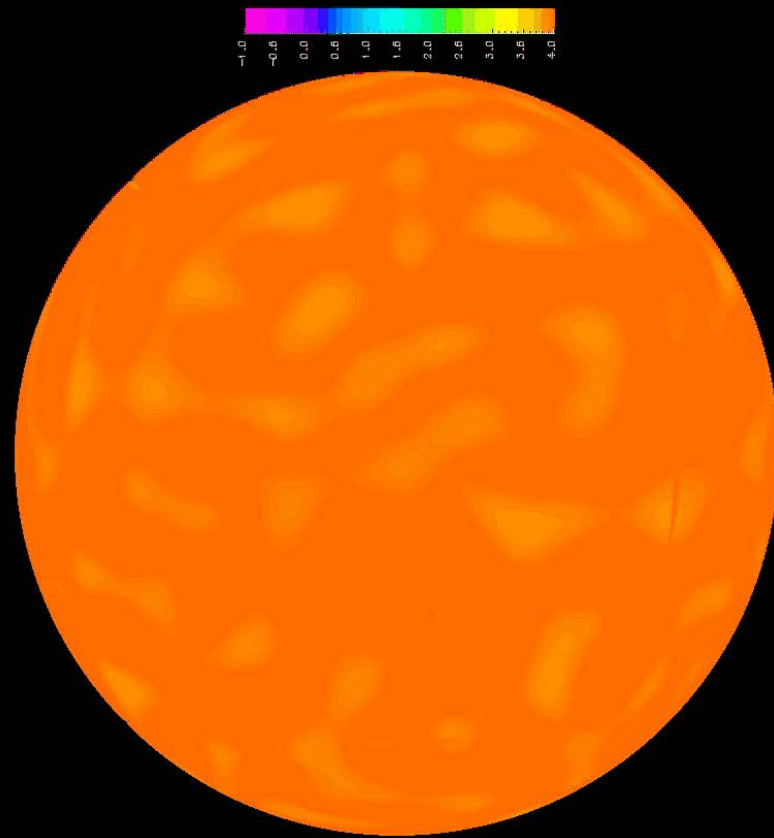




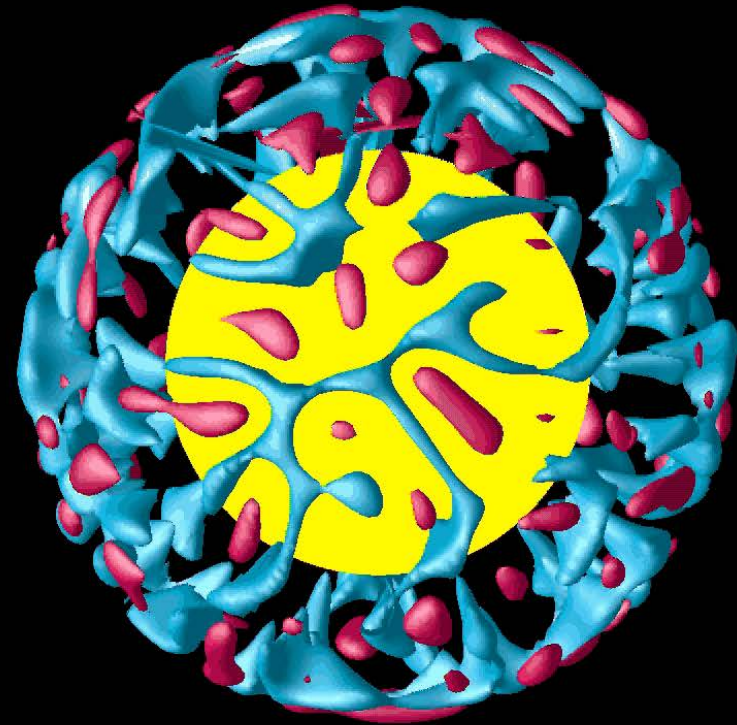
# Stagnant lid convection



# Stagnant lid convection



Yield Stress =  $3.5 \cdot 10000$  (420 MPa)



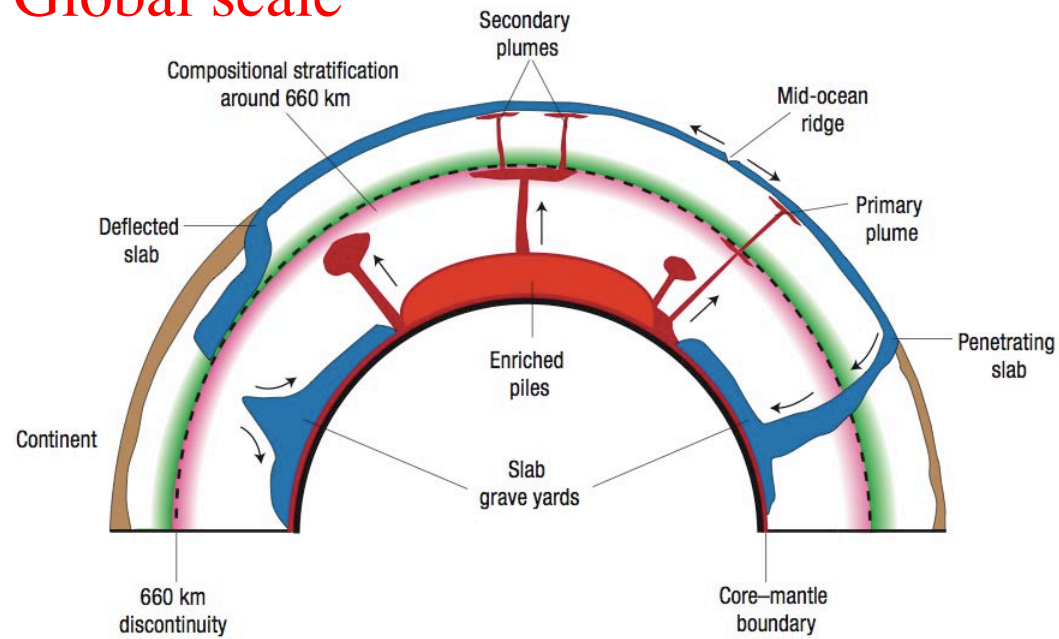
# Why we don't understand plate tectonics at a fundamental level

- Rock deformation is complex
  - Viscous, brittle, plastic, elastic, nonlinear
  - Dependent on grain size, composition (major and trace elements, eg water)
- Multi-scale
  - Lengthscales from mm to 1000s km
  - Timescales from seconds - Gyr



# Multi-scale problem

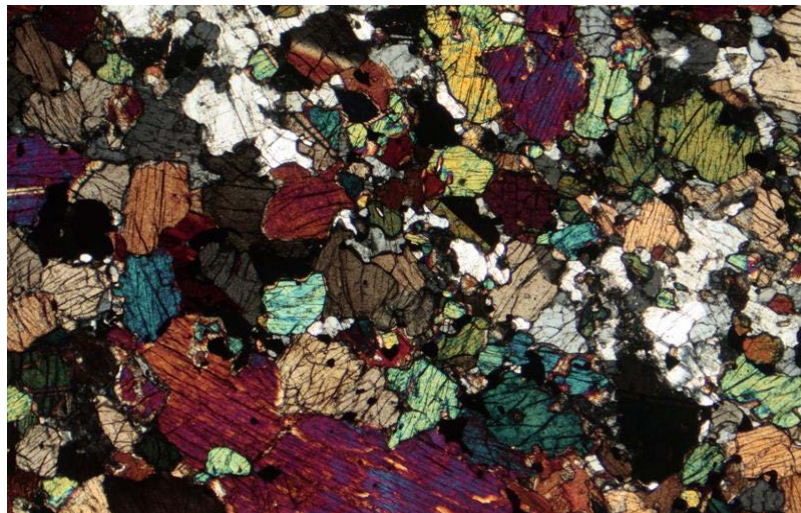
Global scale



'Human' scale



Micro scale

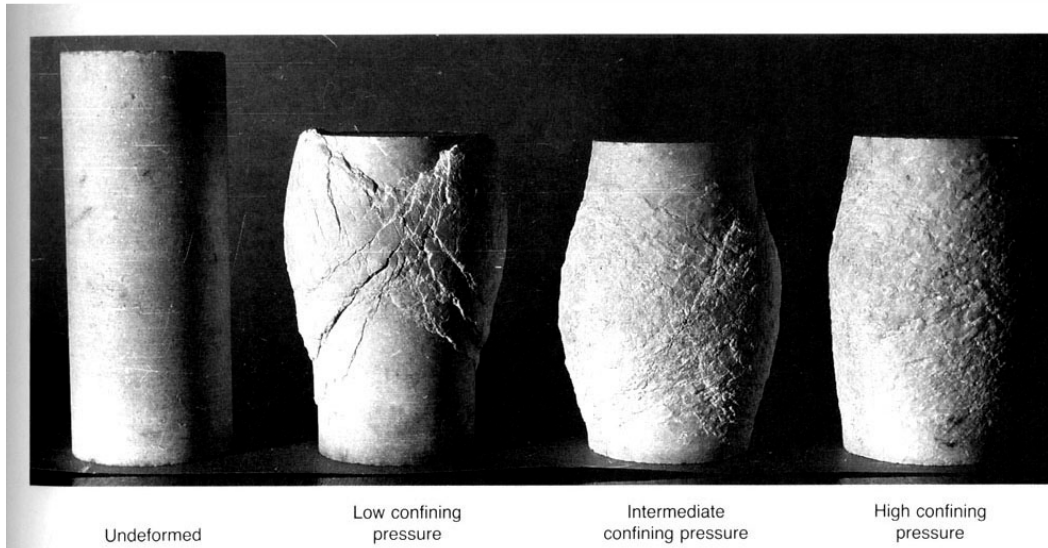




# Strength of rocks

- Increases with confining pressure (depth) then saturates

Low-T deformation: Effect of P



## Low T: Effect of P

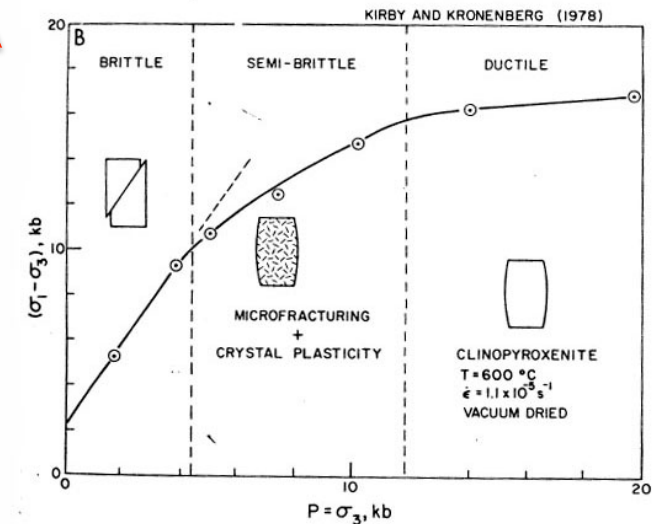
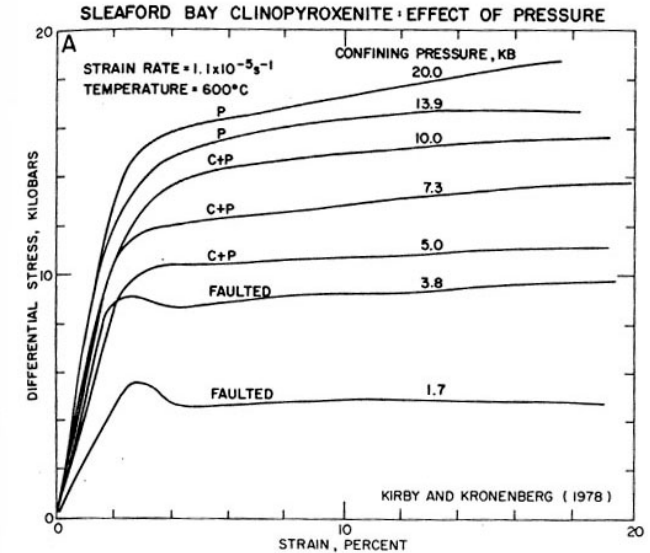
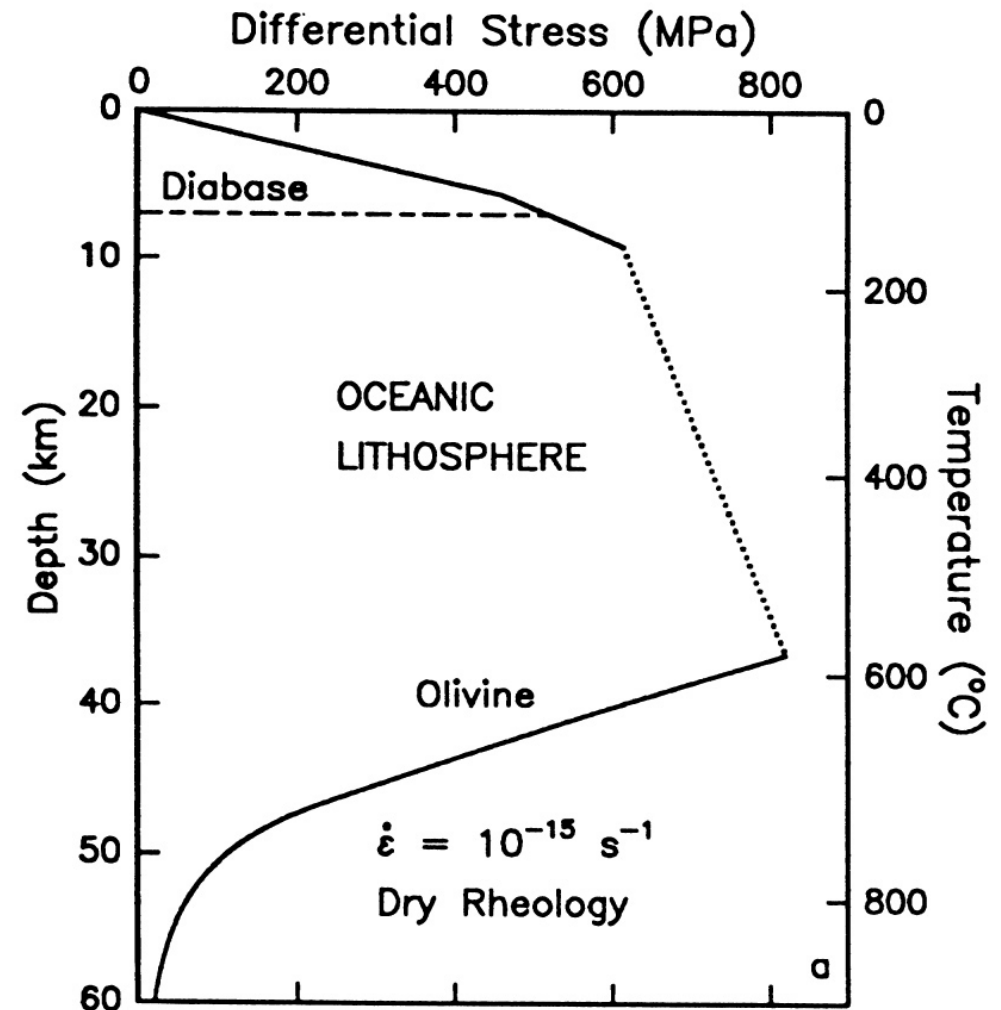
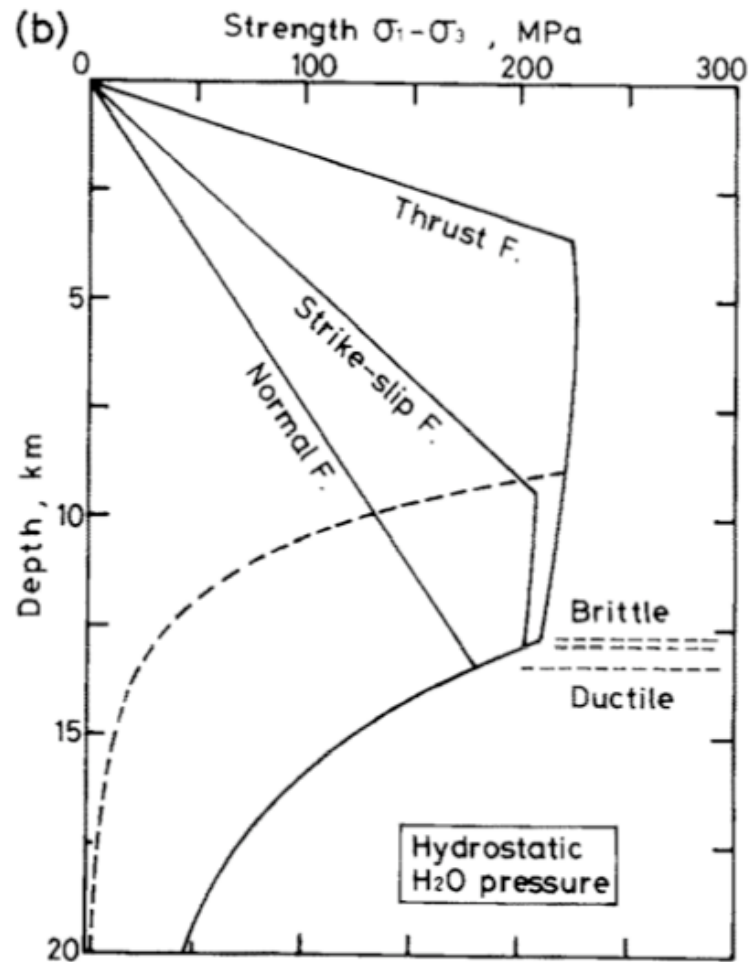


Fig. 6. Effect of confining pressure on the strength of Sleaford Bay clinopyroxenite tested in triaxial compression (S. H. Kirby and A. K. Kronenberg, unpublished data, 1978): (a) stress-strain curves, (b) ultimate strength or stress at 10% strain as a function of confining pressure.

# Strength profile of lithosphere

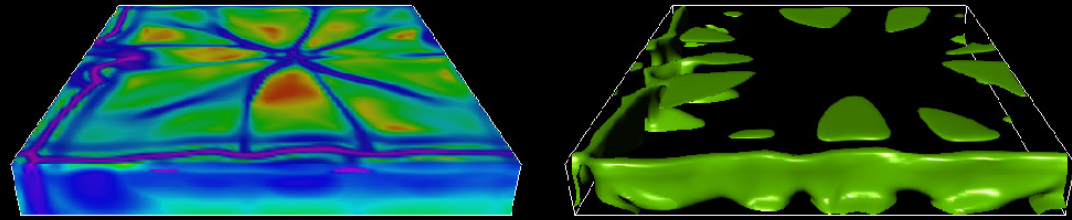
Continental (granite): Shimada 1993

Oceanic: Kohlstedt 1995

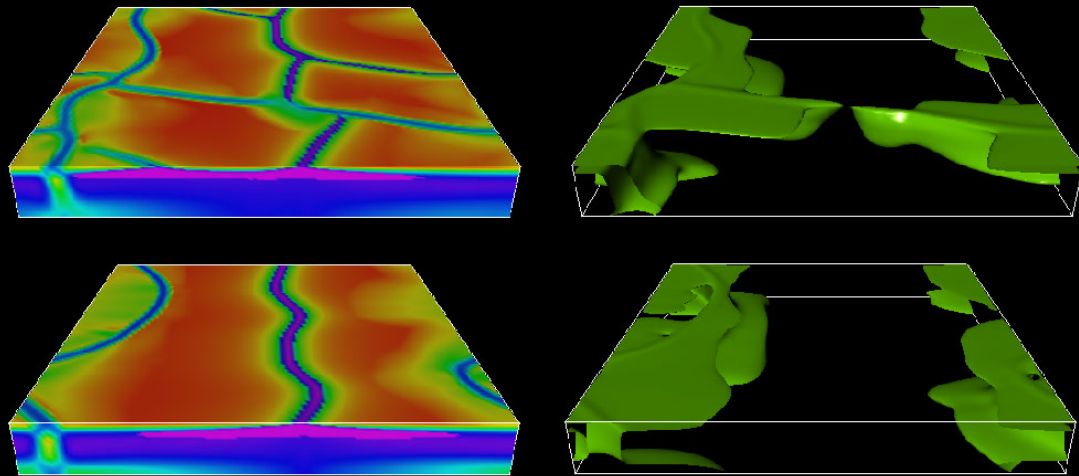


- Varying yield strength, including asthenosph.

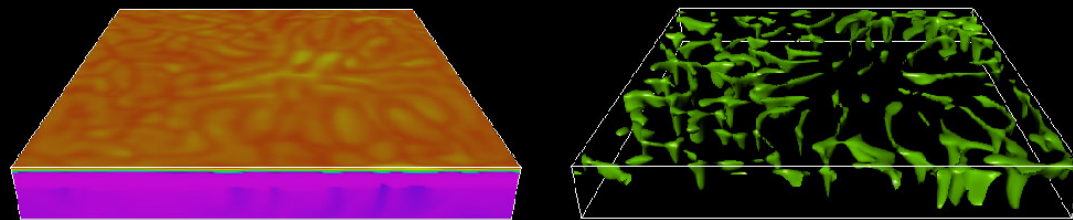
Low yield stress: weak plates, diffuse deformation



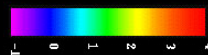
Intermediate yield stress: Good plate tectonics



High yield stress: Immobile lithosphere



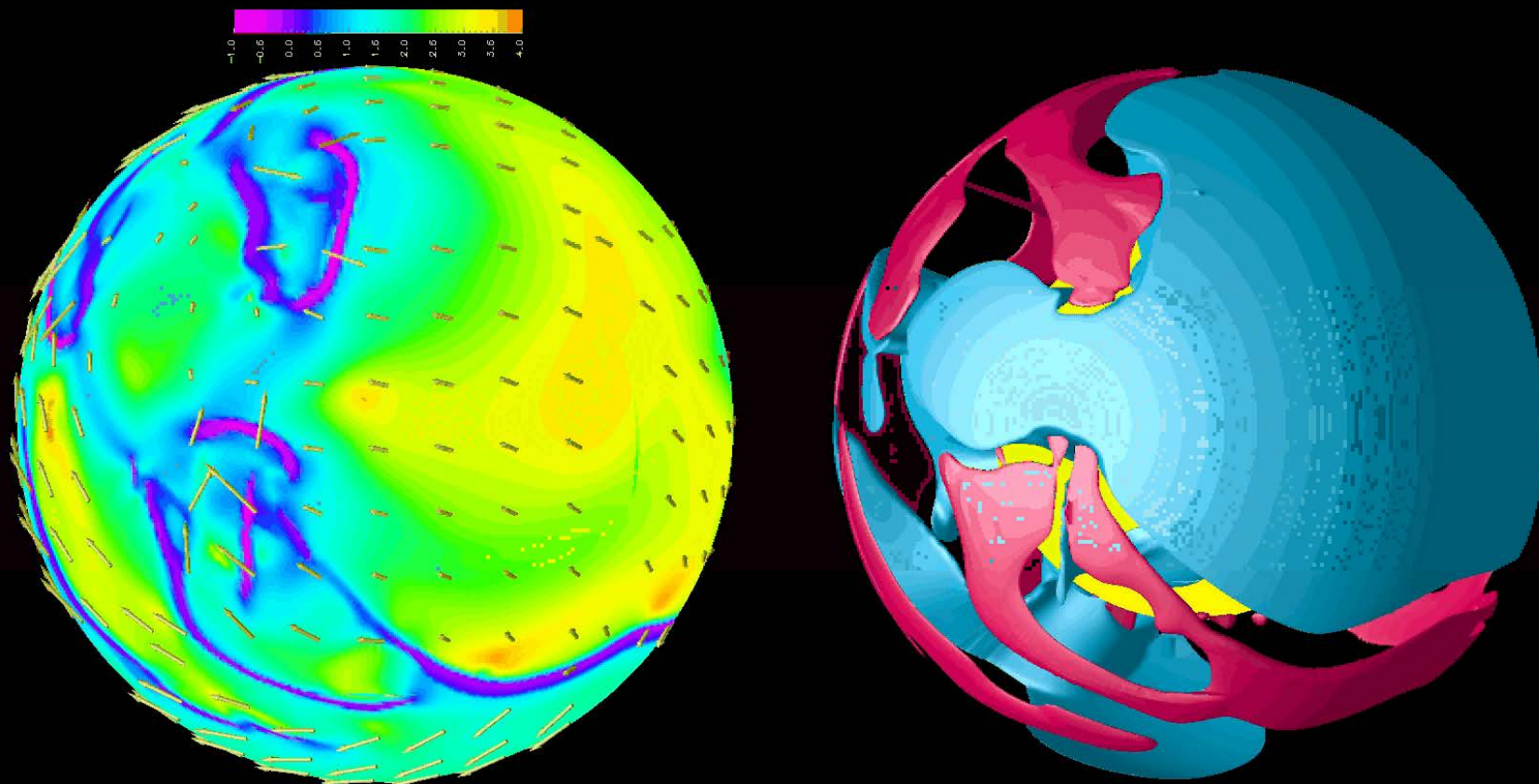
viscosity



cold T (downwellings)

by Paul J. Tackley 2000

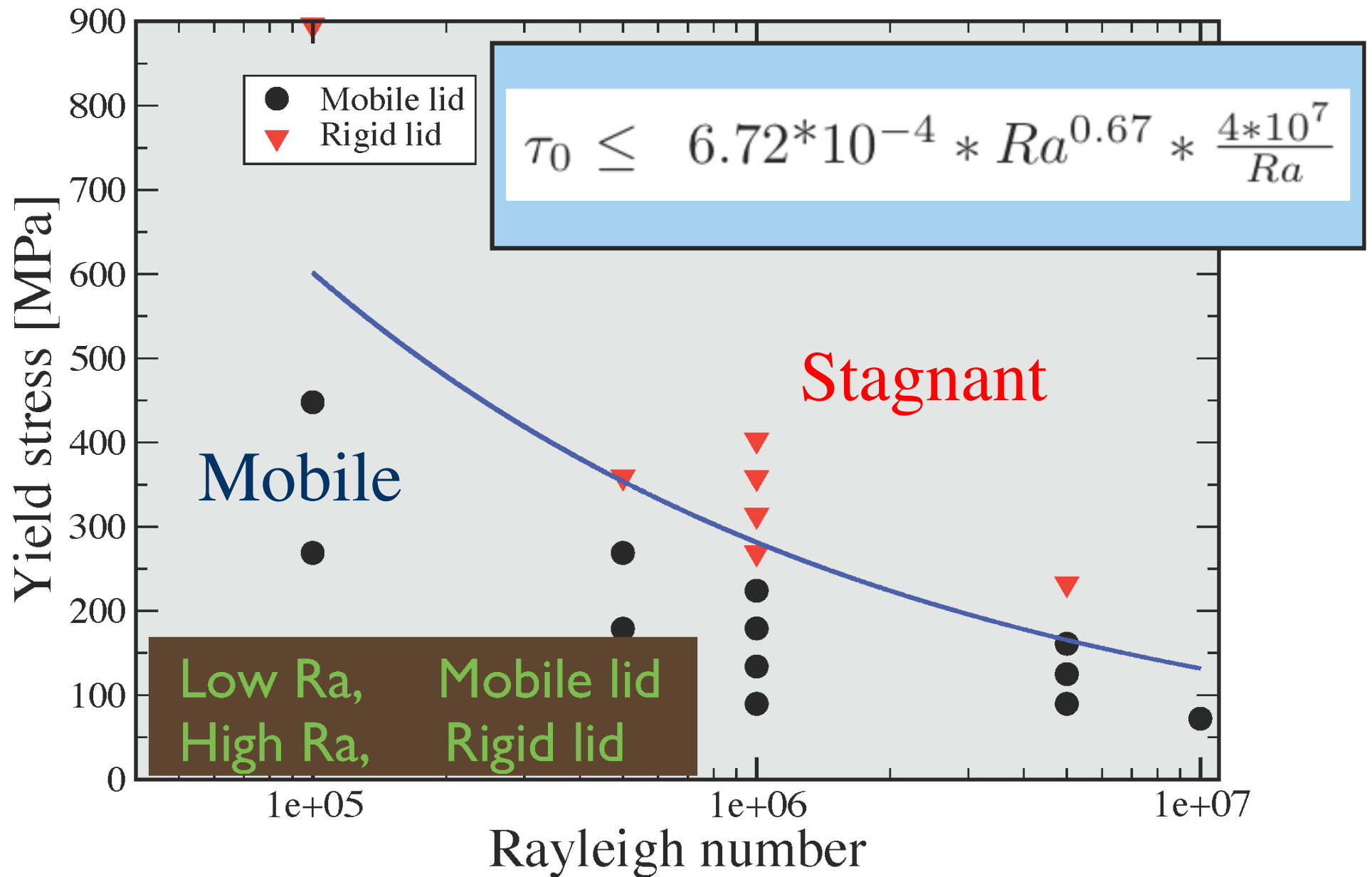
# Mobile lid mode



H. Van Heck

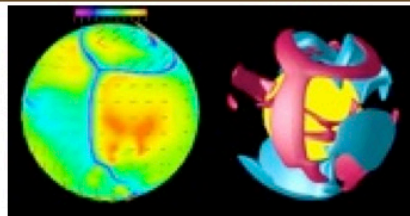


# Rayleigh number versus yield stress

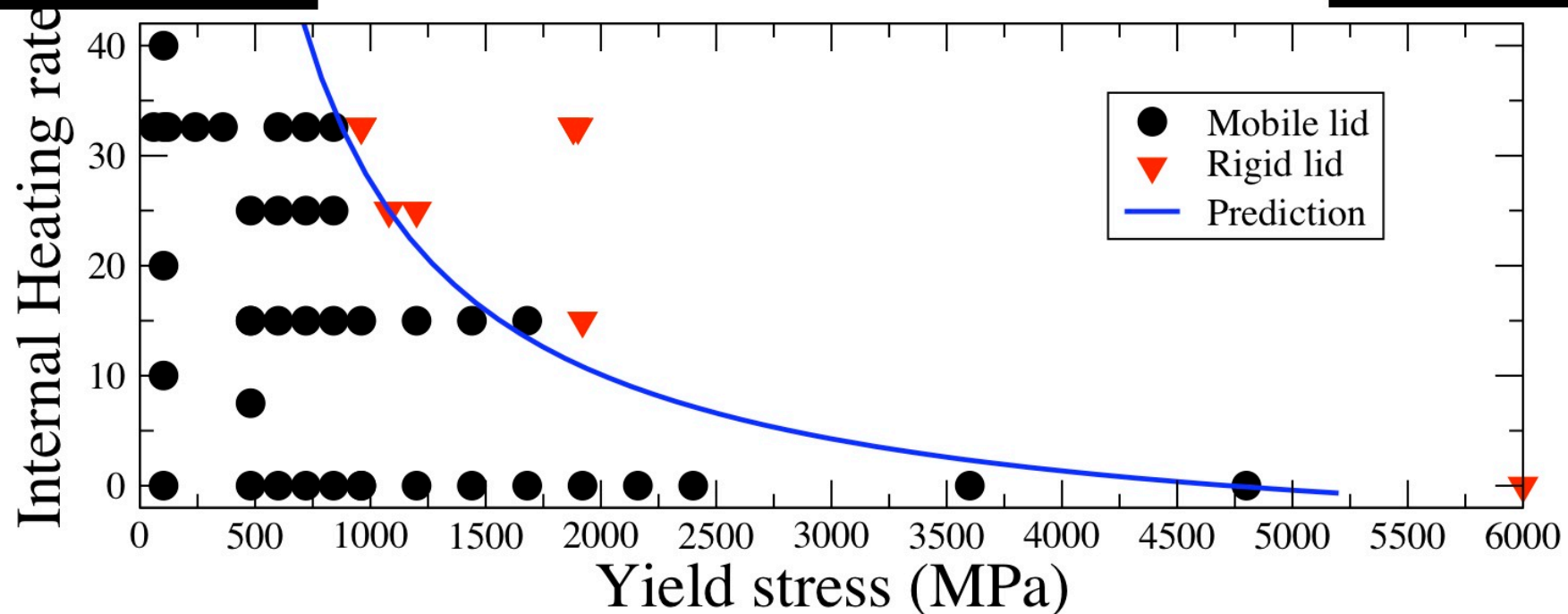
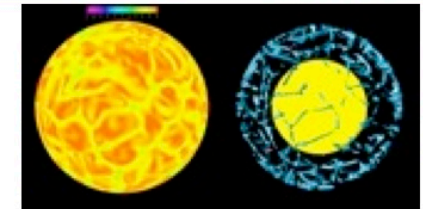


# Internal heating rate

## Strength of the lithosphere vs convective stresses



Yield stress versus internal heating rate



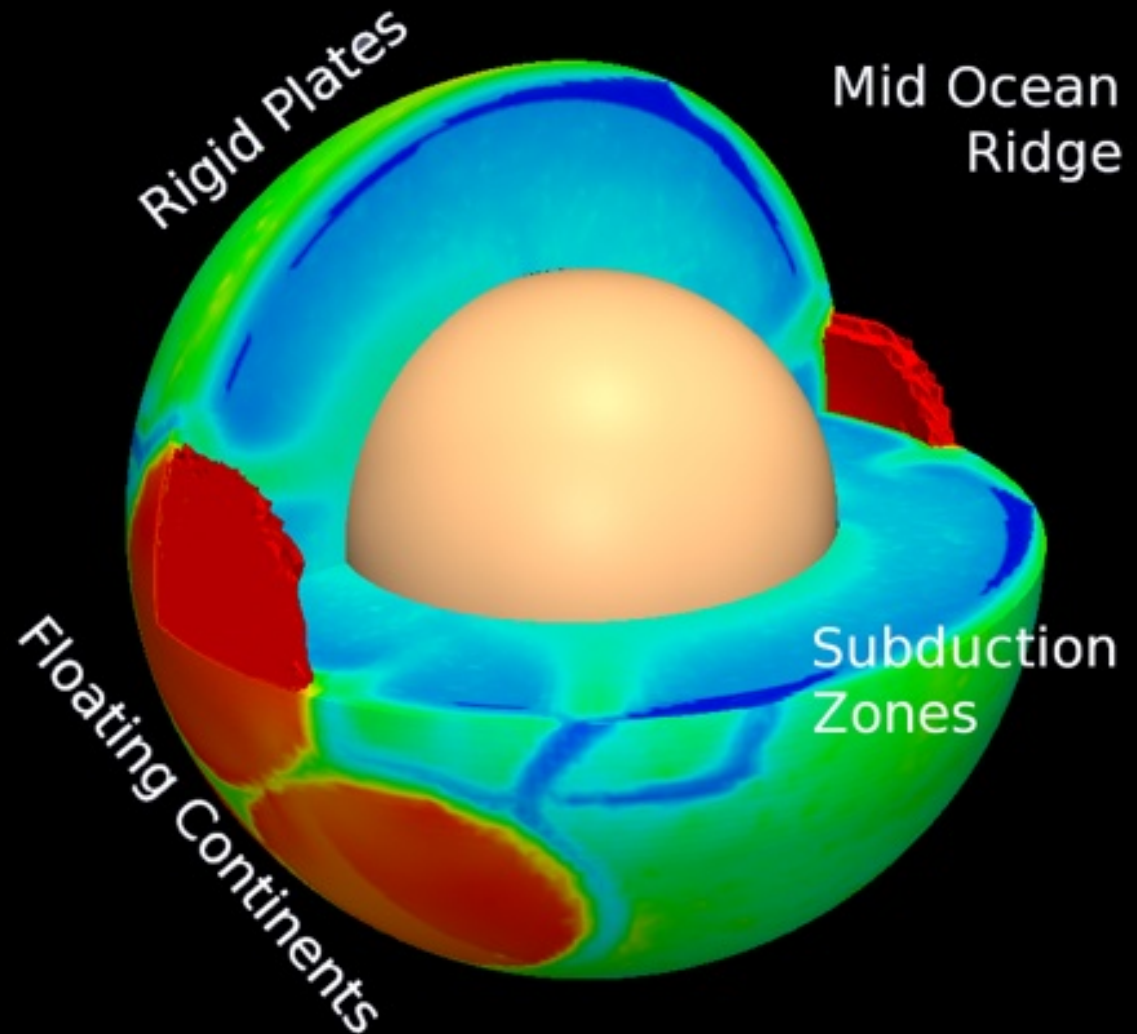
# Implications for terrestrial planet evolution

- Plate tectonics favoured at
  - higher mantle viscosity (lower Ra)
  - Lower internal (radioactive) heating
- Both predict transition **stagnant lid->plates** as planet cools.

# Influence of **continents** on self-consistent plate tectonics?



Tobias Rolf



**CRYSTAL2PLATE**  
How does plate tectonics work:  
From crystal-scale processes to mantle  
convection with self-consistent plates



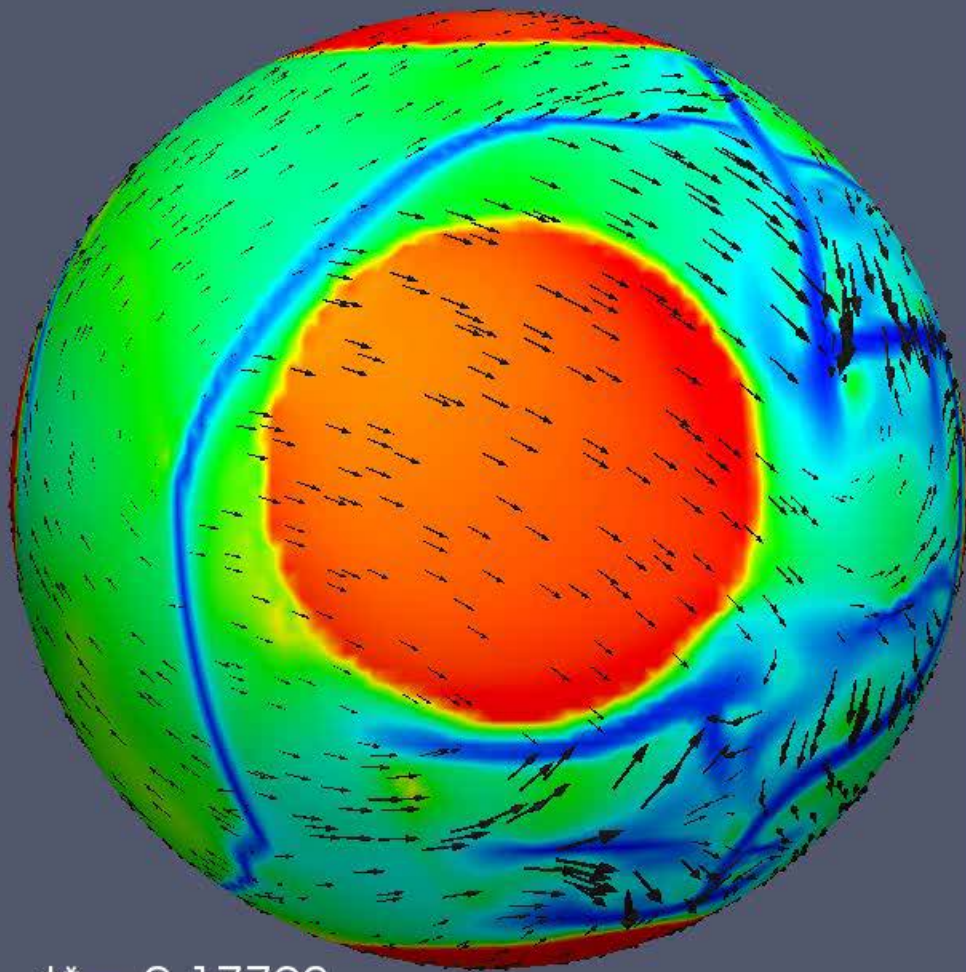
MARIE CURIE ACTIONS



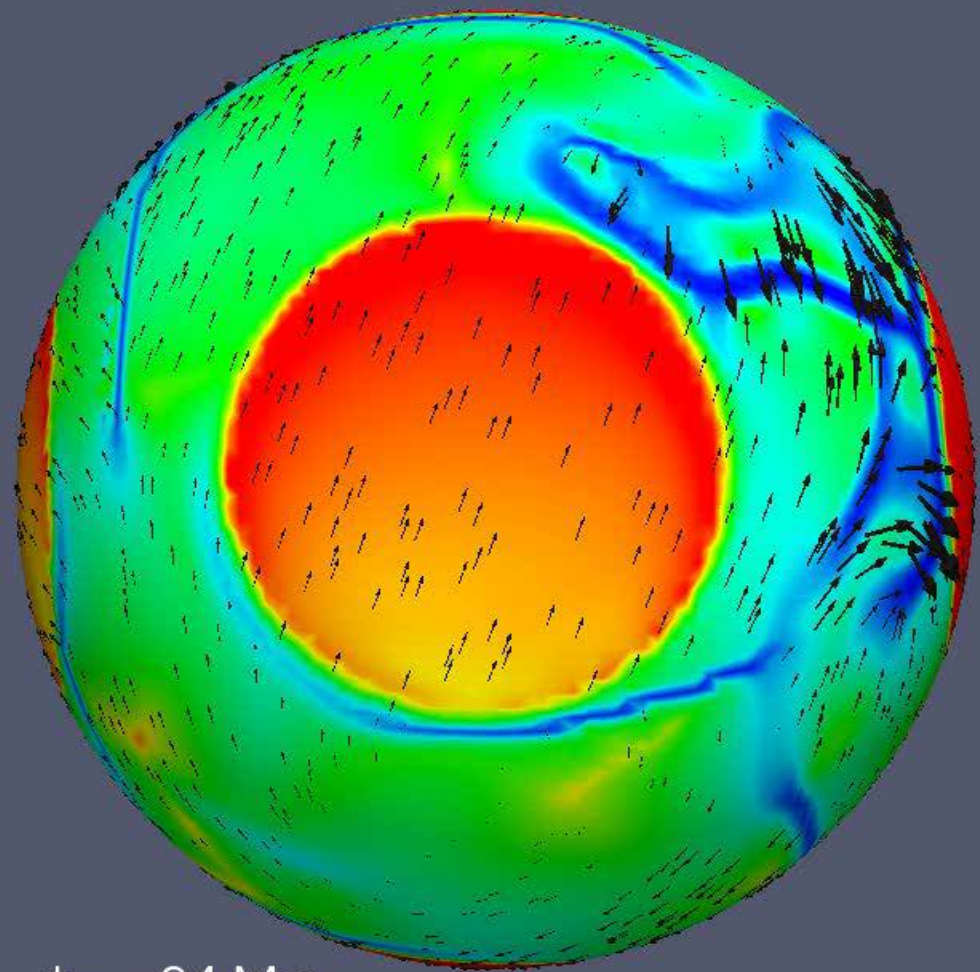
SEVENTH FRAMEWORK  
PROGRAMME



0.5 100000



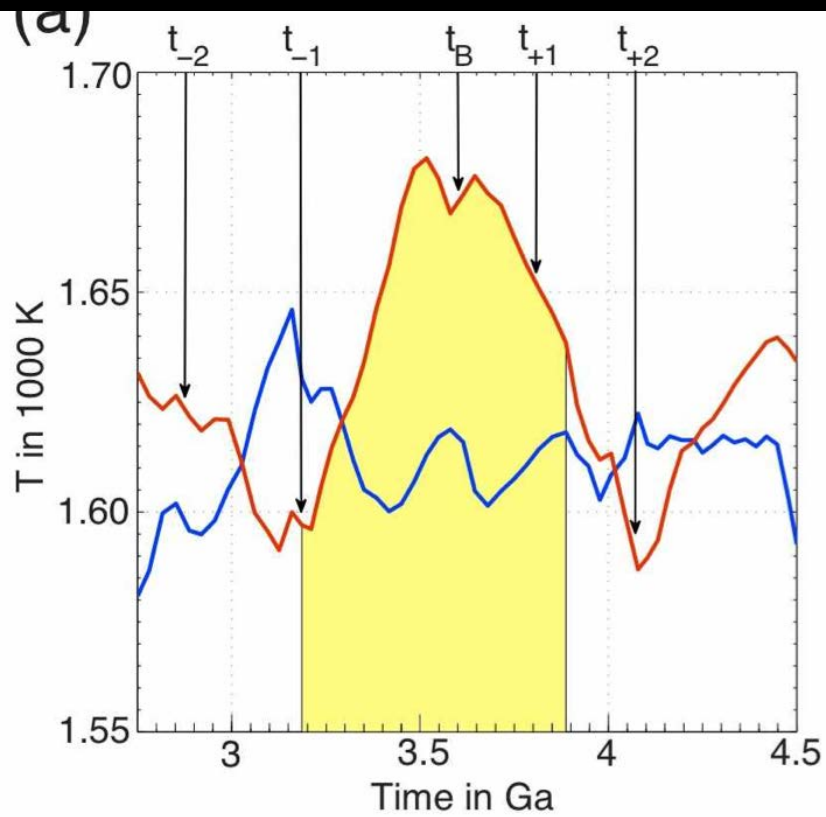
$t^* = 0.17792$



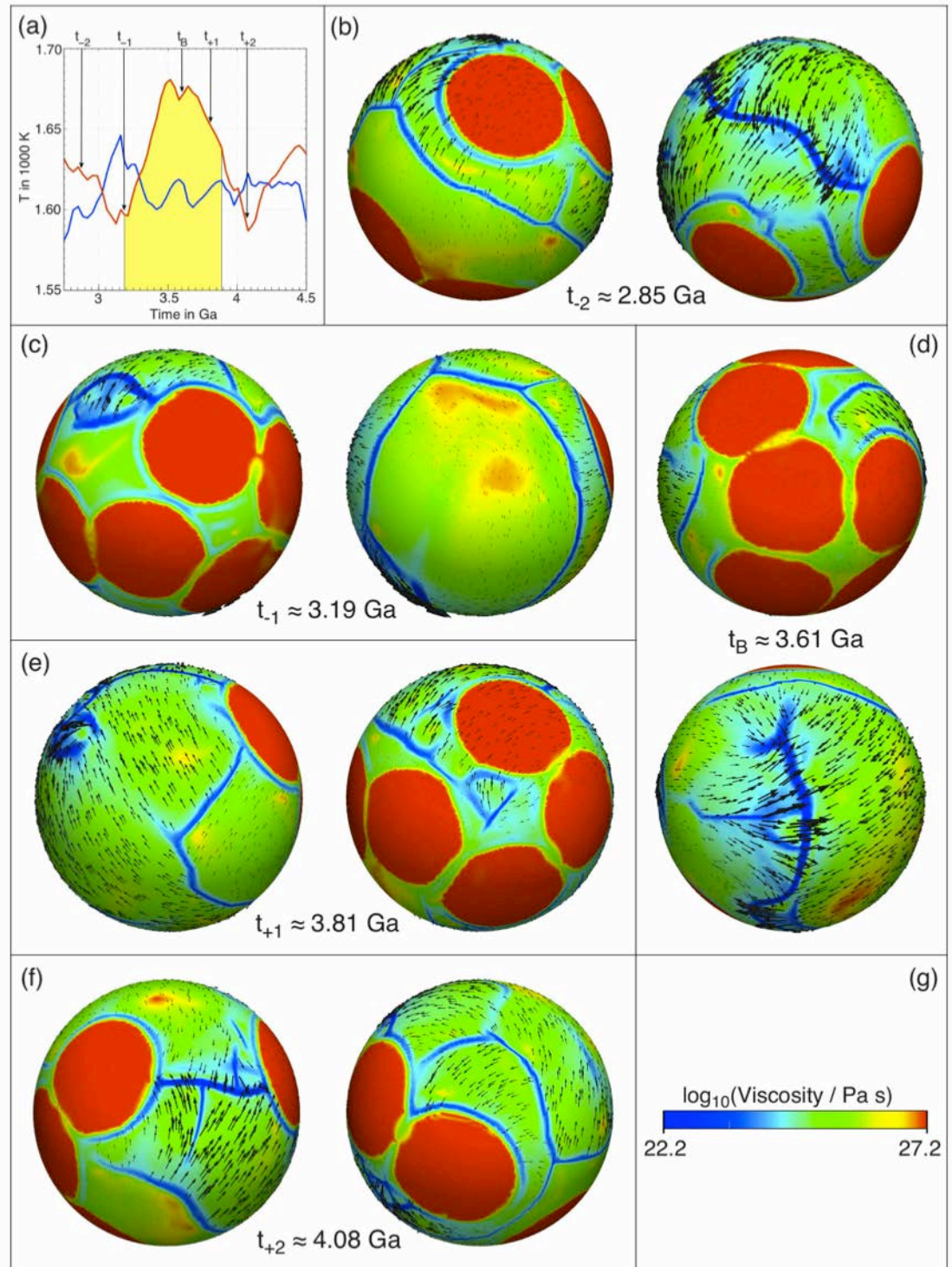
$t = -24 \text{ Ma}$

6 continents

# Supercontinent cycle



Warming under  
supercontinent

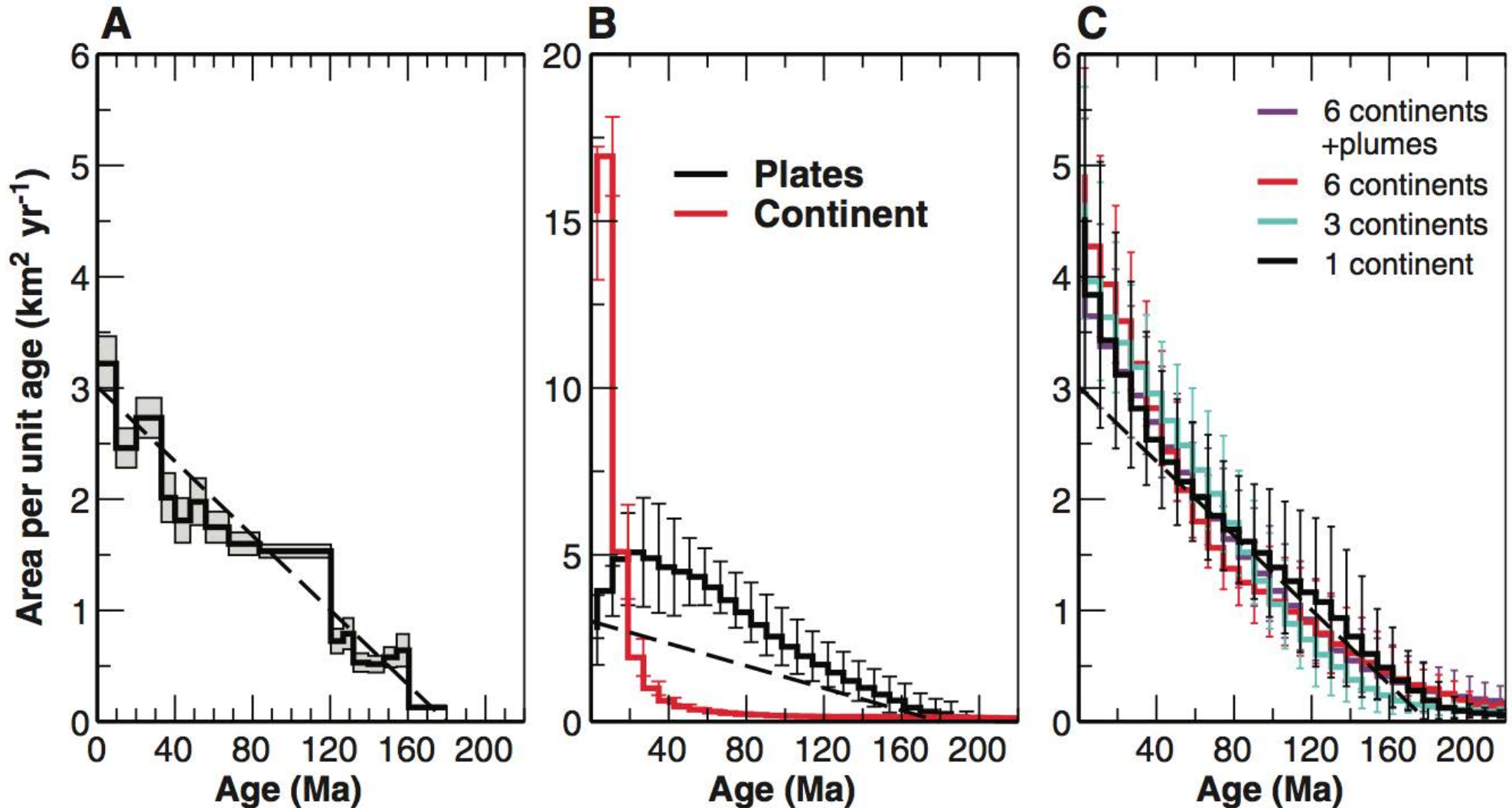


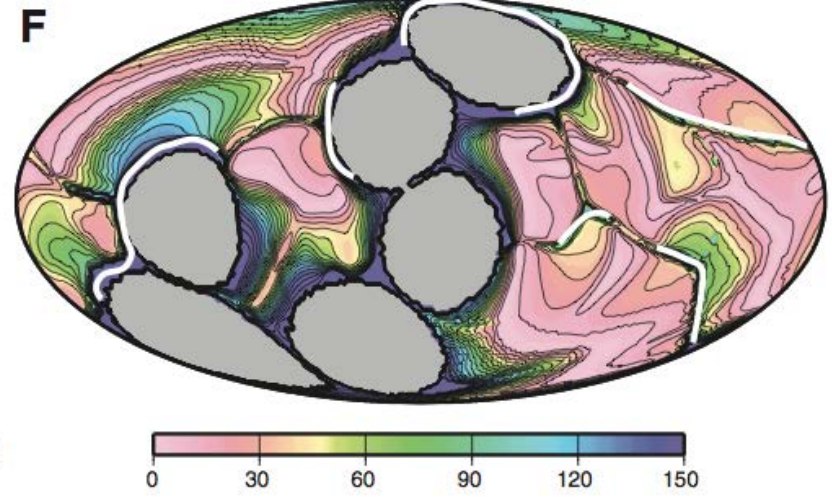
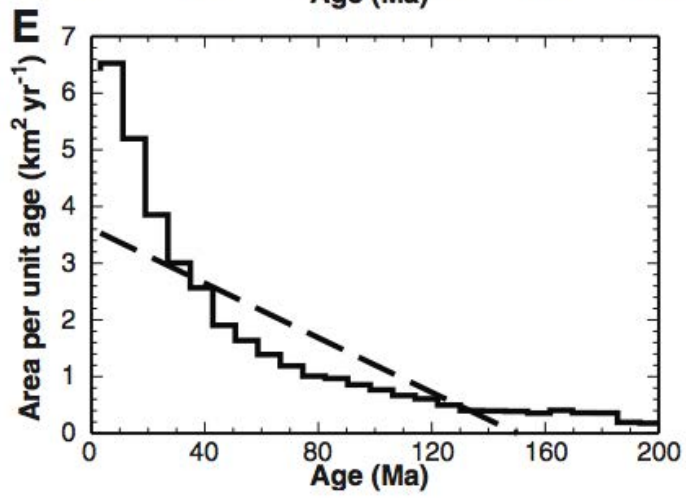
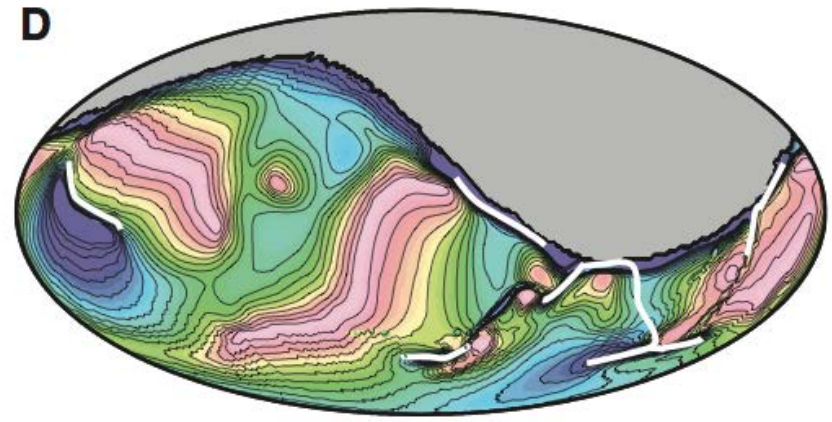
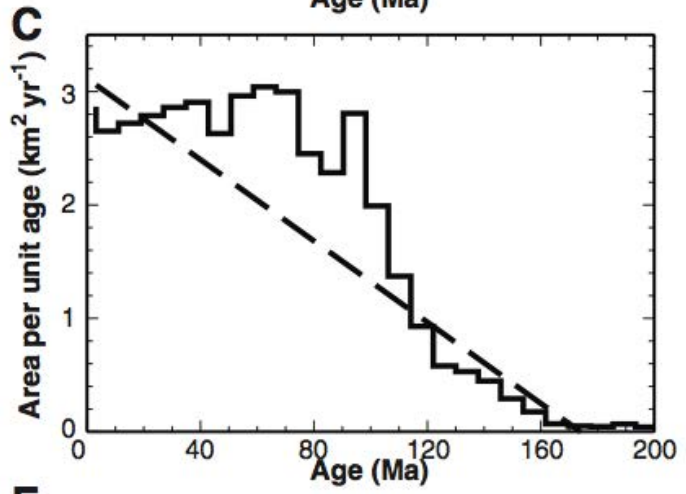
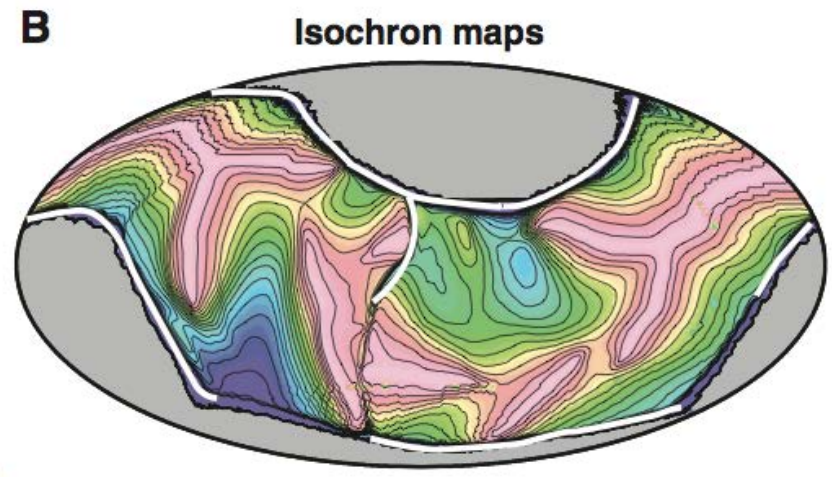
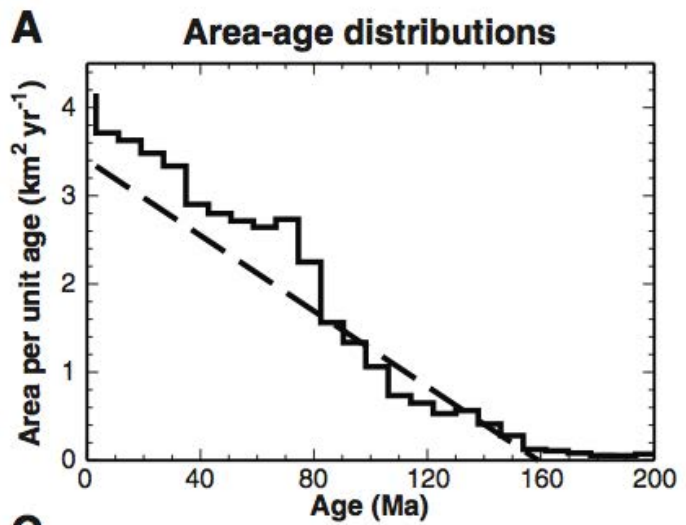


# Dynamic Causes of the Relation Between Area and Age of the Ocean Floor

N. Coltice,<sup>1,2\*</sup> T. Rolf,<sup>3</sup> P. J. Tackley,<sup>3</sup> S. Labrosse<sup>1,2</sup>

SCIENCE VOL 336 20 APRIL 2012





Distribution shape varies with time



# Reason 1: geometry of plate boundaries

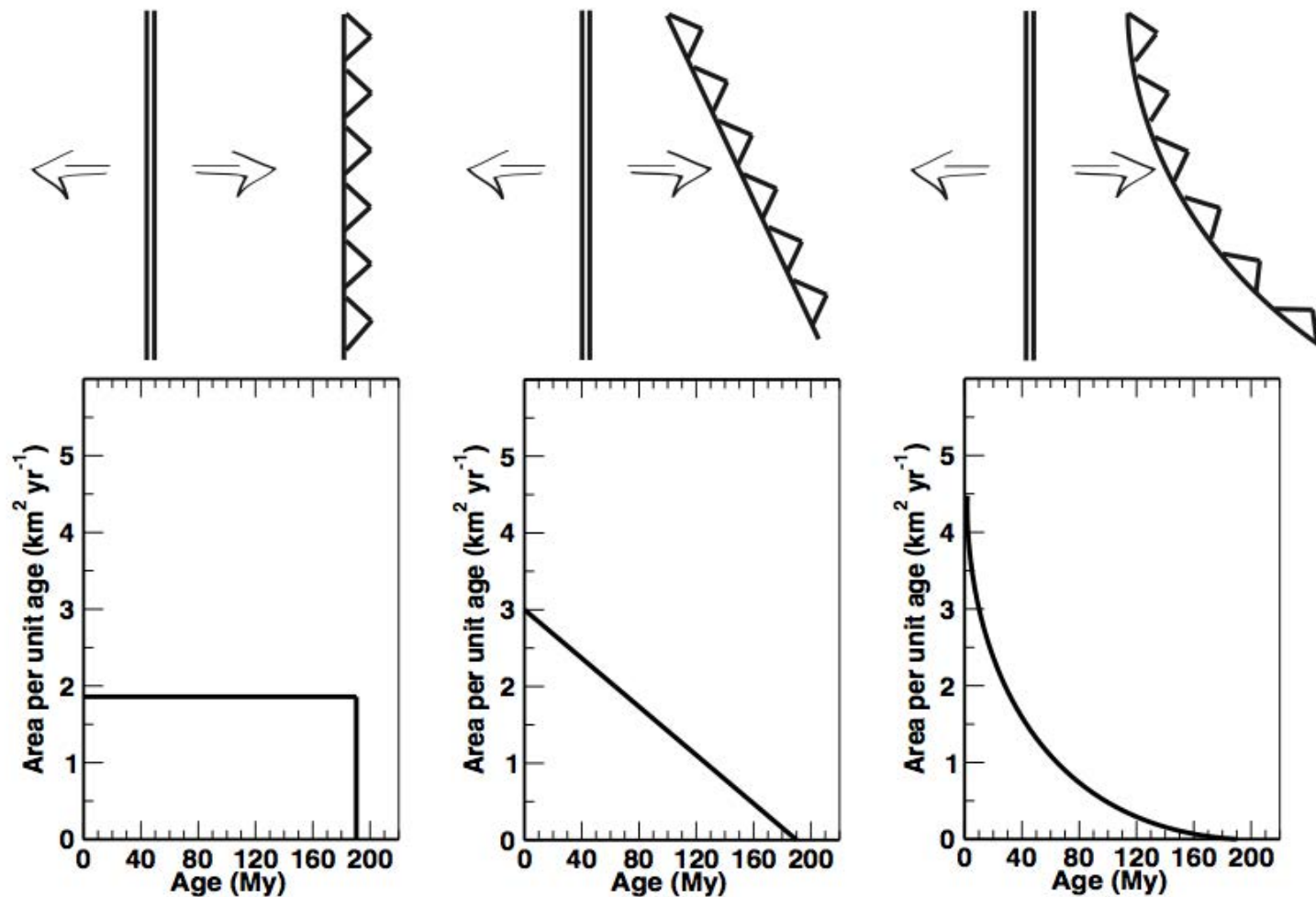


Figure 1: The shape of the area-age distribution depends on the geometry of plate boundaries.

## Reason 2: time-dependence of seafloor production

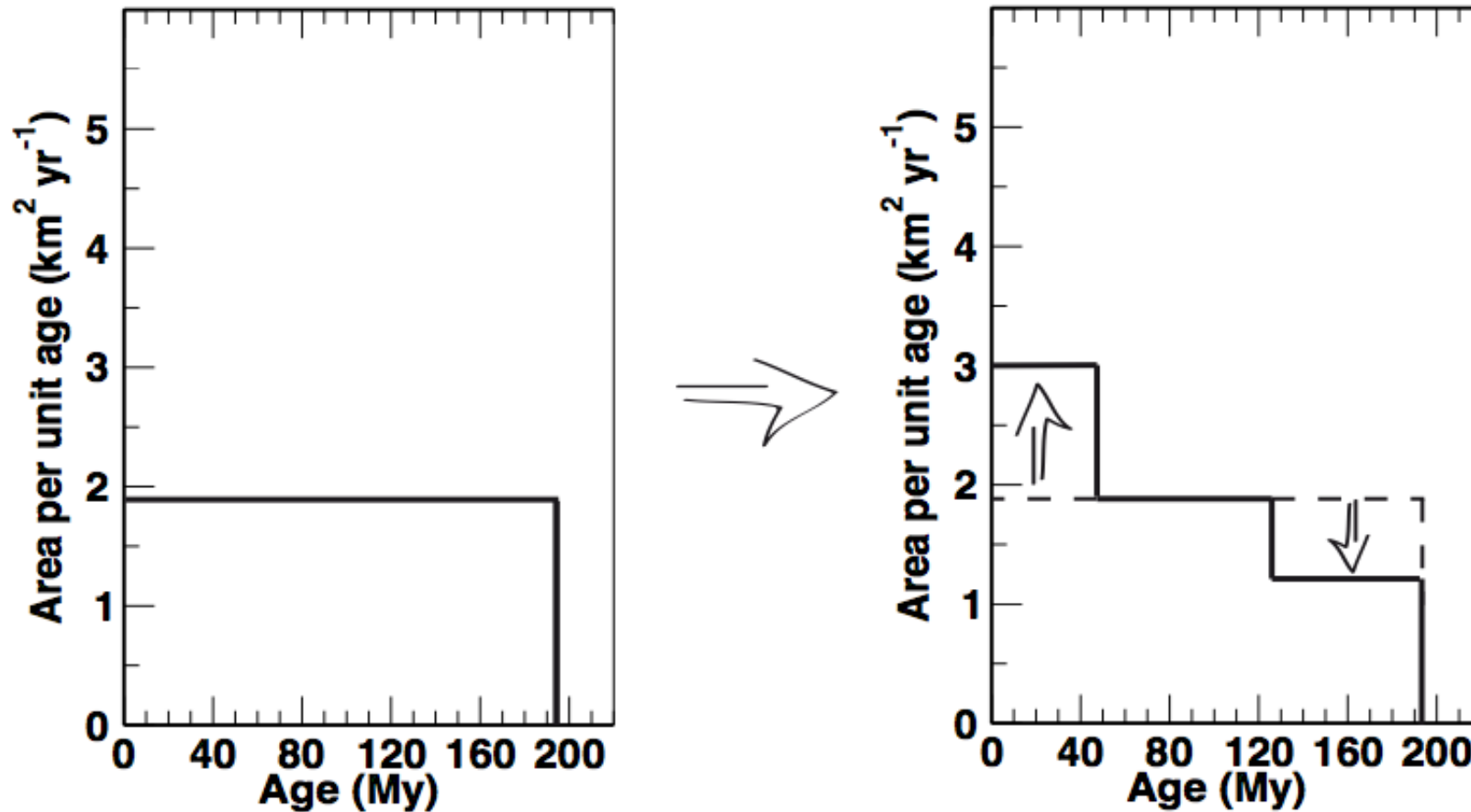
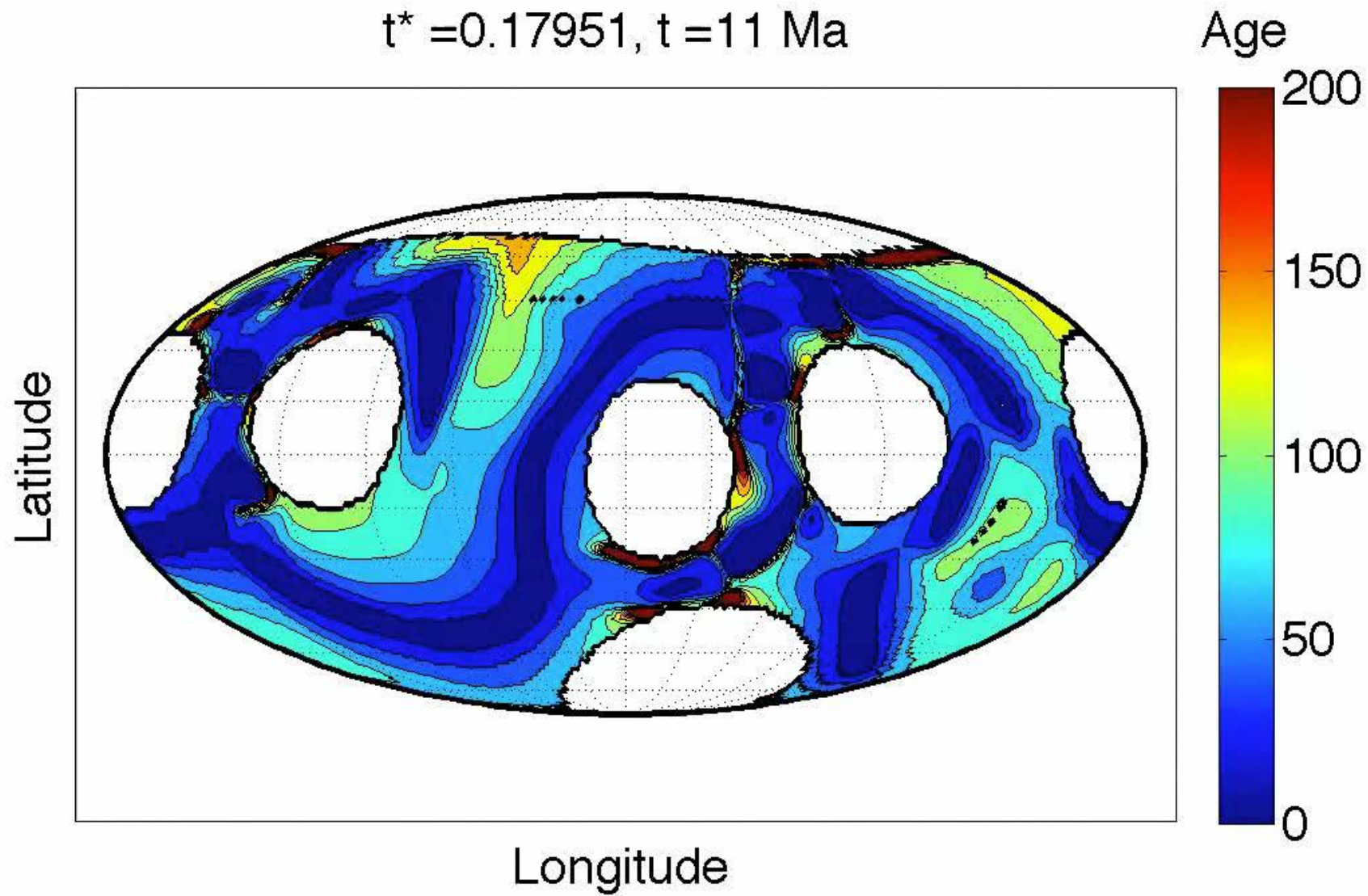


Figure 2: The shape of the area-age distribution depends on the time dependence of the production of new seafloor.

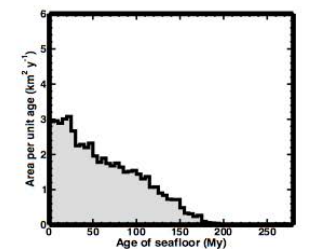
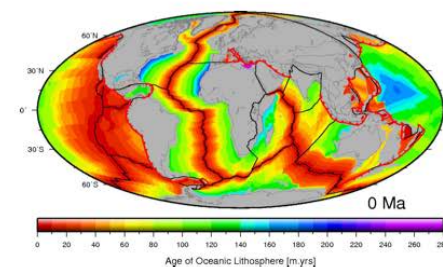
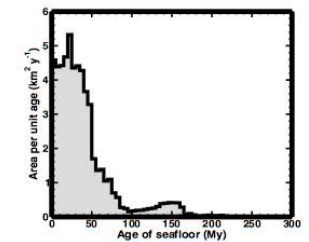
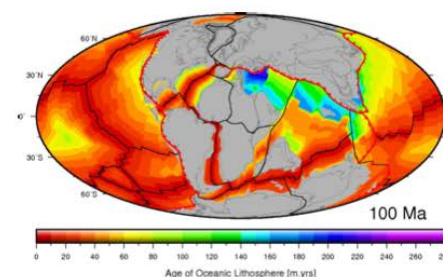
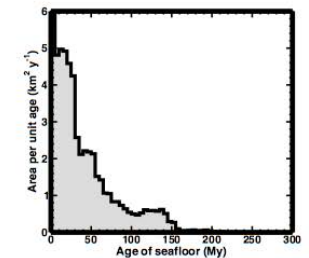
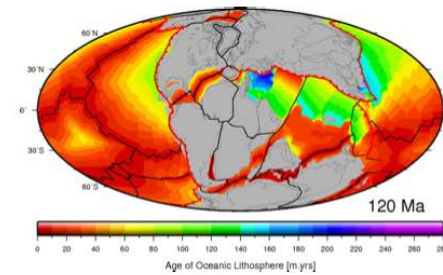
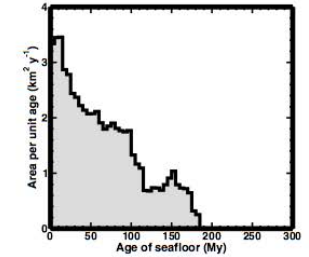
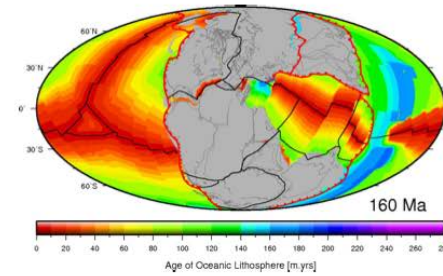
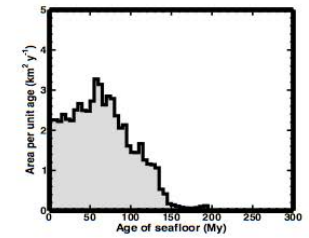
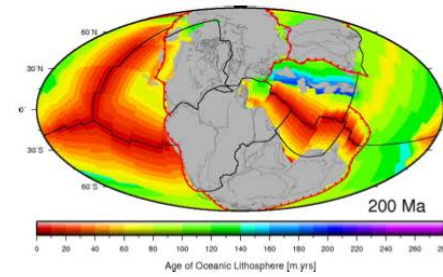
$t^* = 0.17951, t = 11 \text{ Ma}$



# The real Earth

Maps of reconstructed distribution of seafloor ages and associated area-age distributions in the past 200 Myr (Seton et al. 2012)

Coltice et al. 2013 EPSL

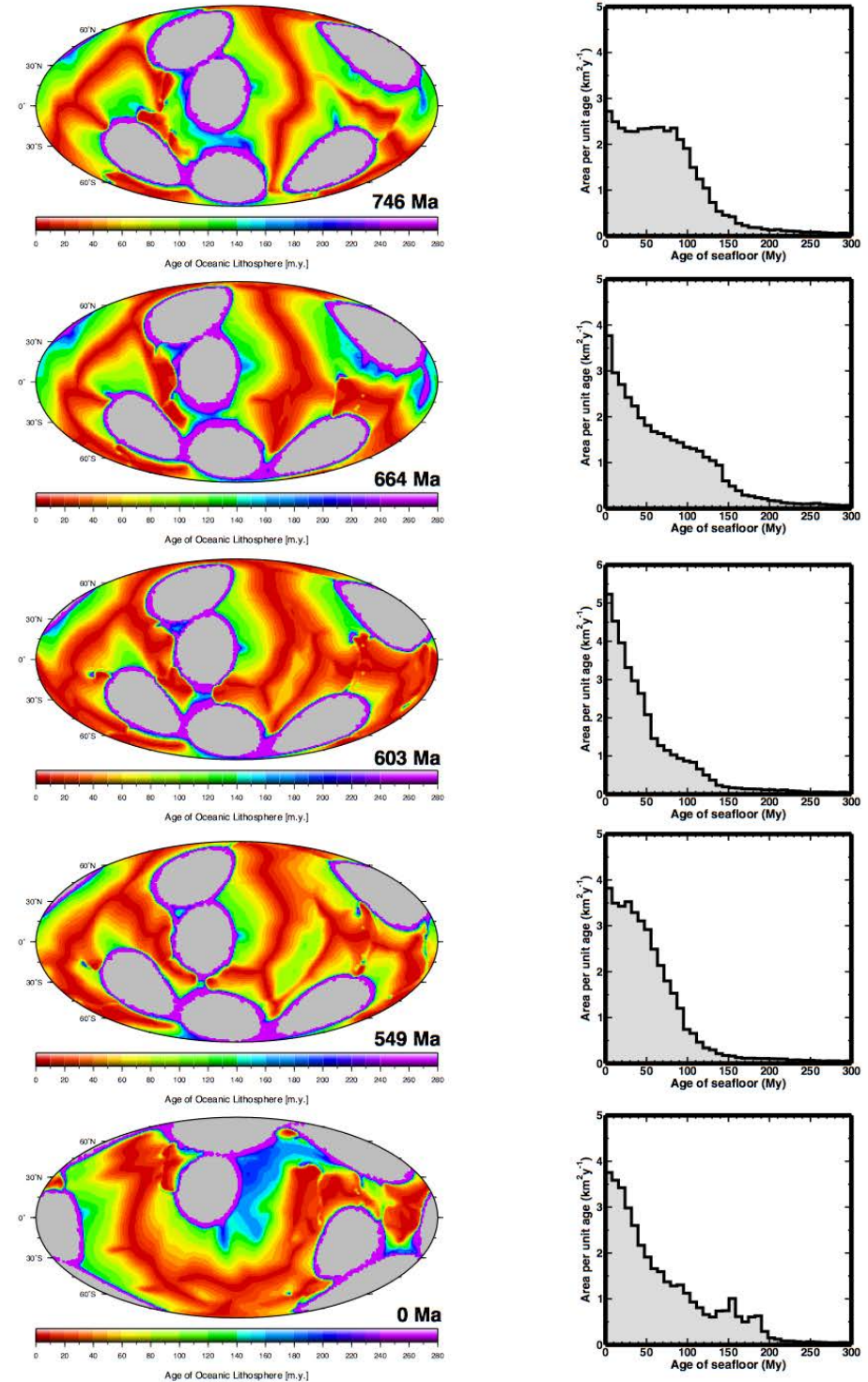




# The simulated Earth

Synthetic maps of seafloor ages and associated area-age distributions in the mantle convection model with 6 continental rafts and 14% of core heating.

Coltice et al. 2013 EPSL



Subduction controls the distribution and fragmentation of Earth's tectonic plates, C. Mallard et al., Nature 2016



# Influence of yield stress on plate distribution



Claire  
Mallard

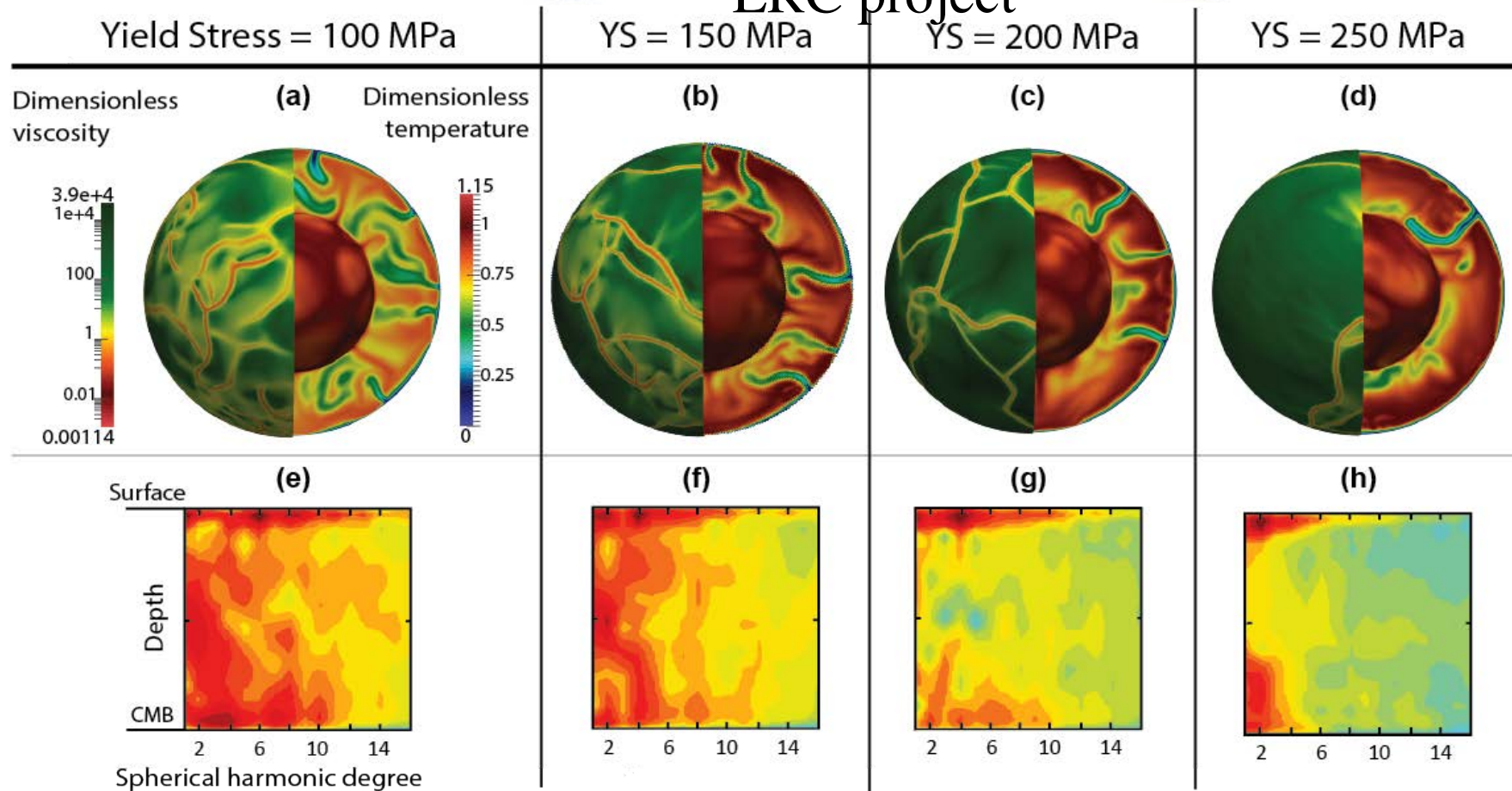


Nicolas  
Coltice  
AUGURY



Dietmar  
Müller

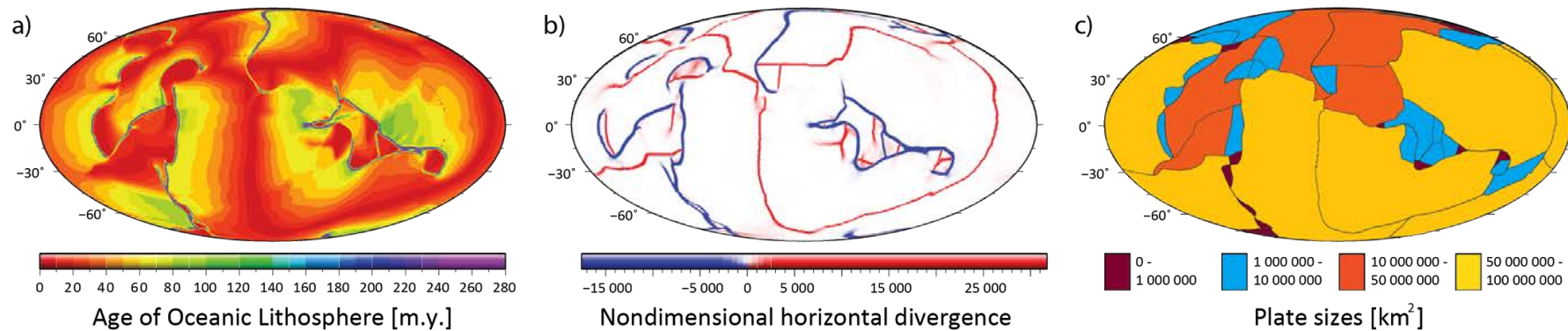
ERC project



Higher yield stress -> larger plates



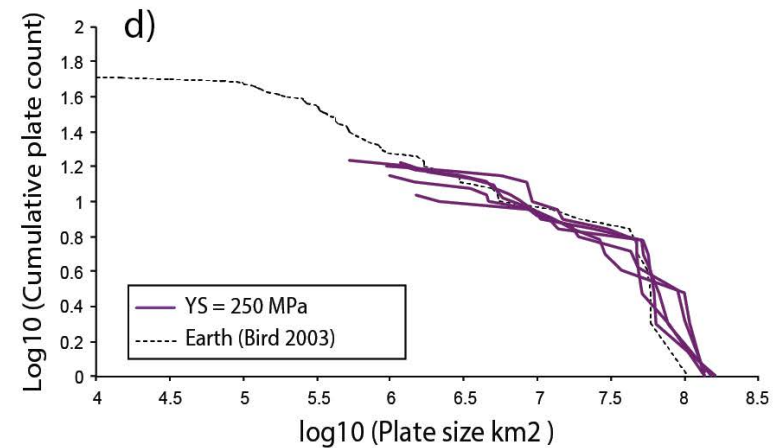
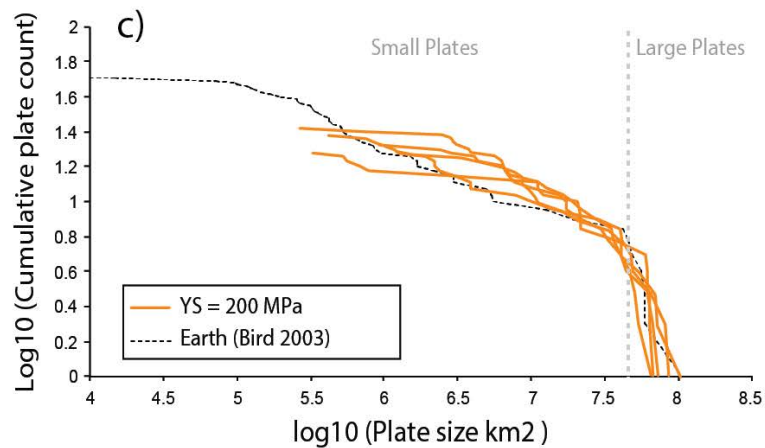
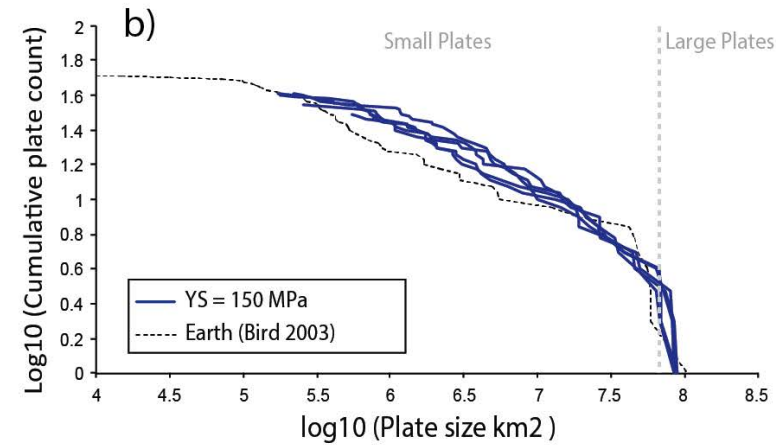
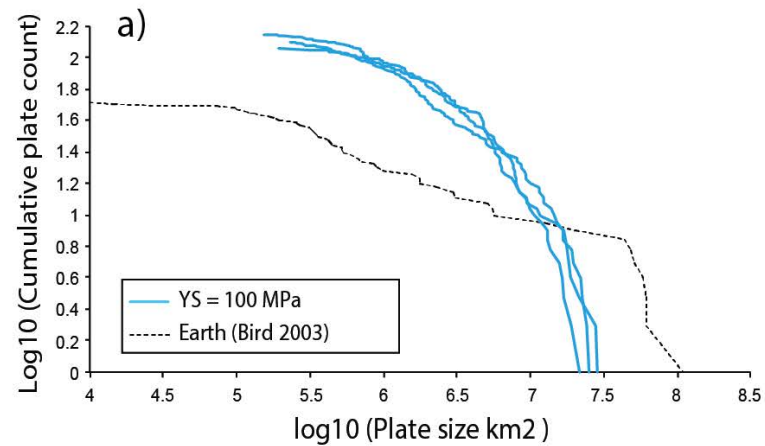
# Generate plate maps



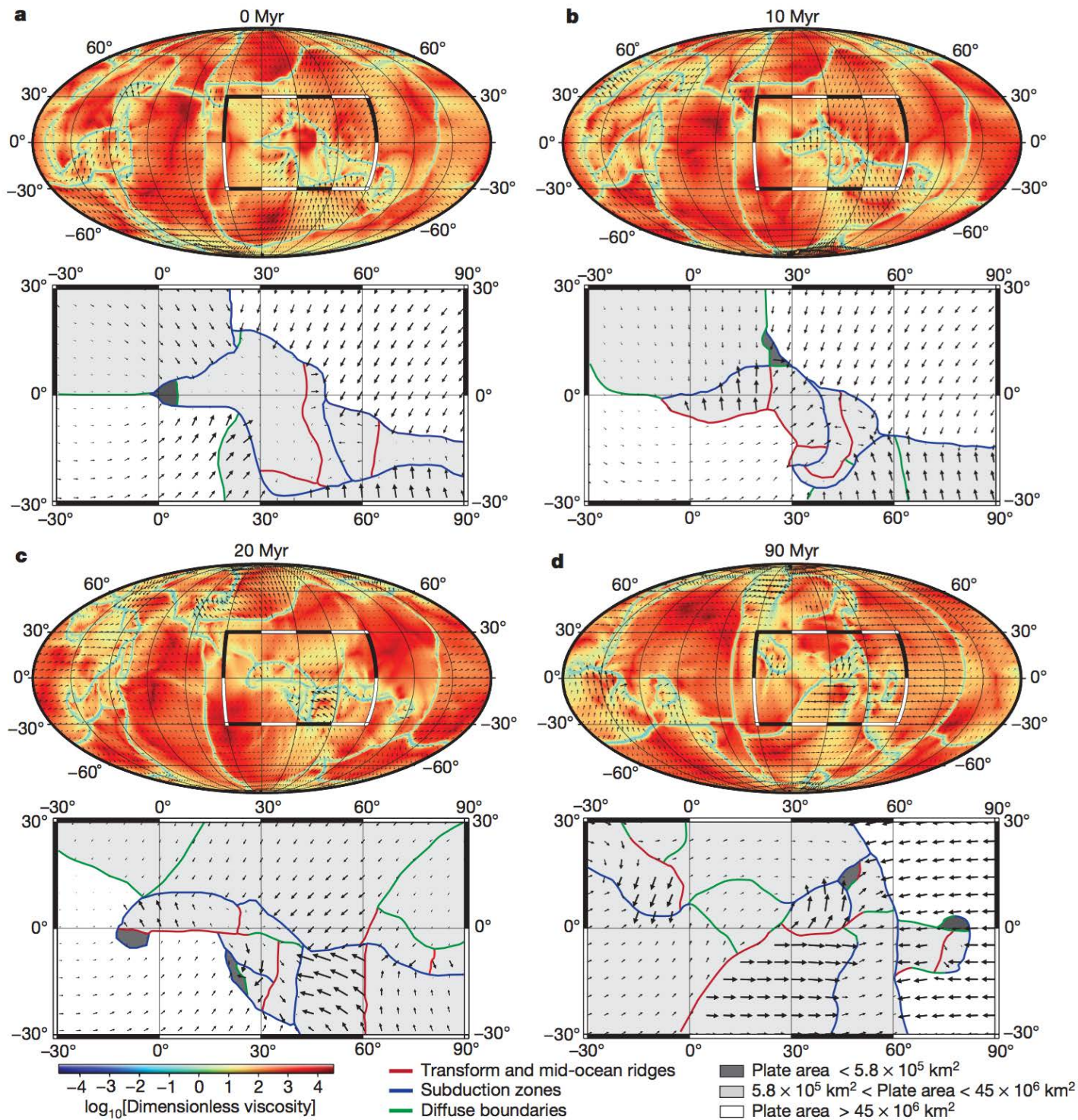
Plates are not a uniform size  
-> a distribution from large to small



# Plate size distribution



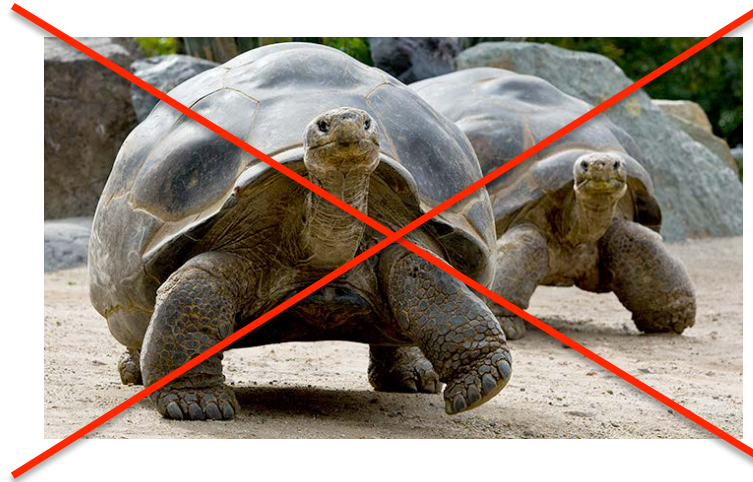
**YS=200 MPa -> matches Earth!**



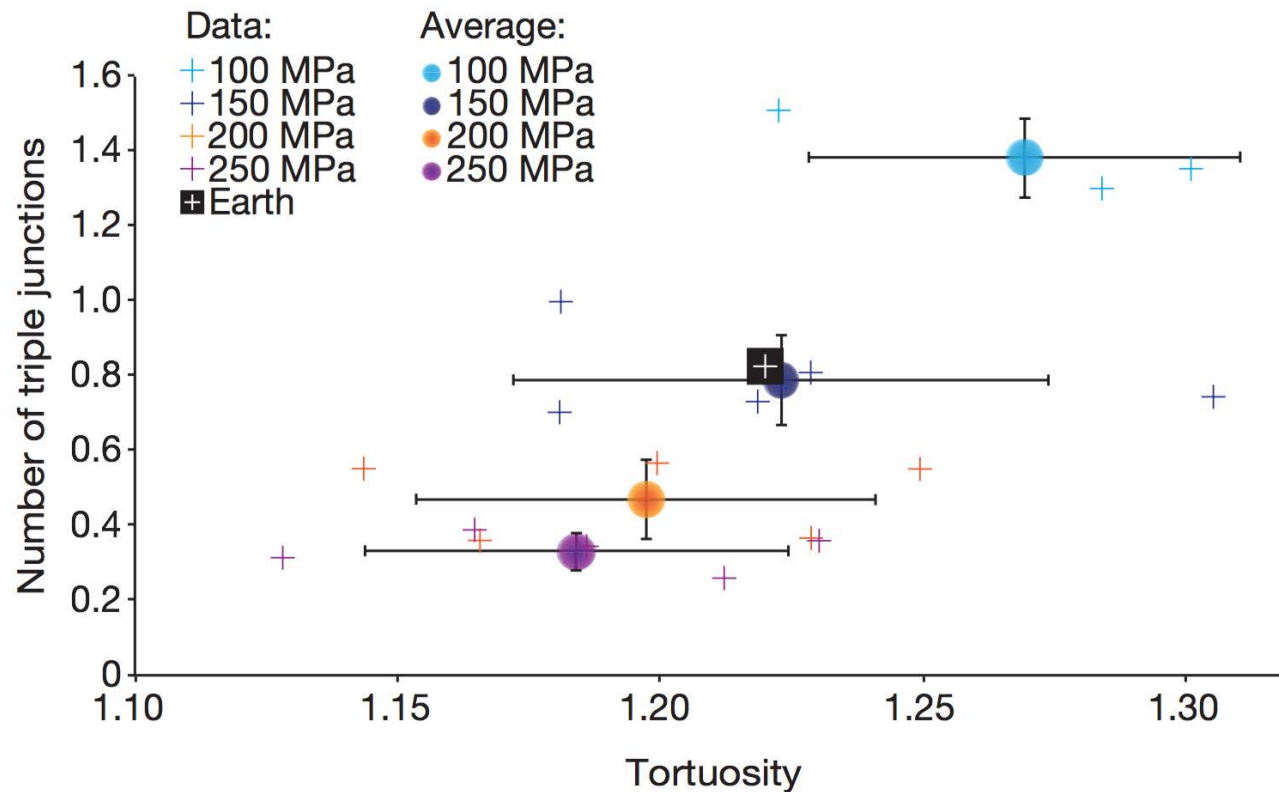
**High curvature  
 => new small  
 plates form**

**Large plates  
 stay ~same**

# Tortuosity!



**More tortuous -> more triple junctions**



**Hence,**

**-> a simple rheology**

**reproduces many features**

**of Earth's plate tectonics**

**(age-area distribution; size distribution) BUT**

**-> required yield stress is**

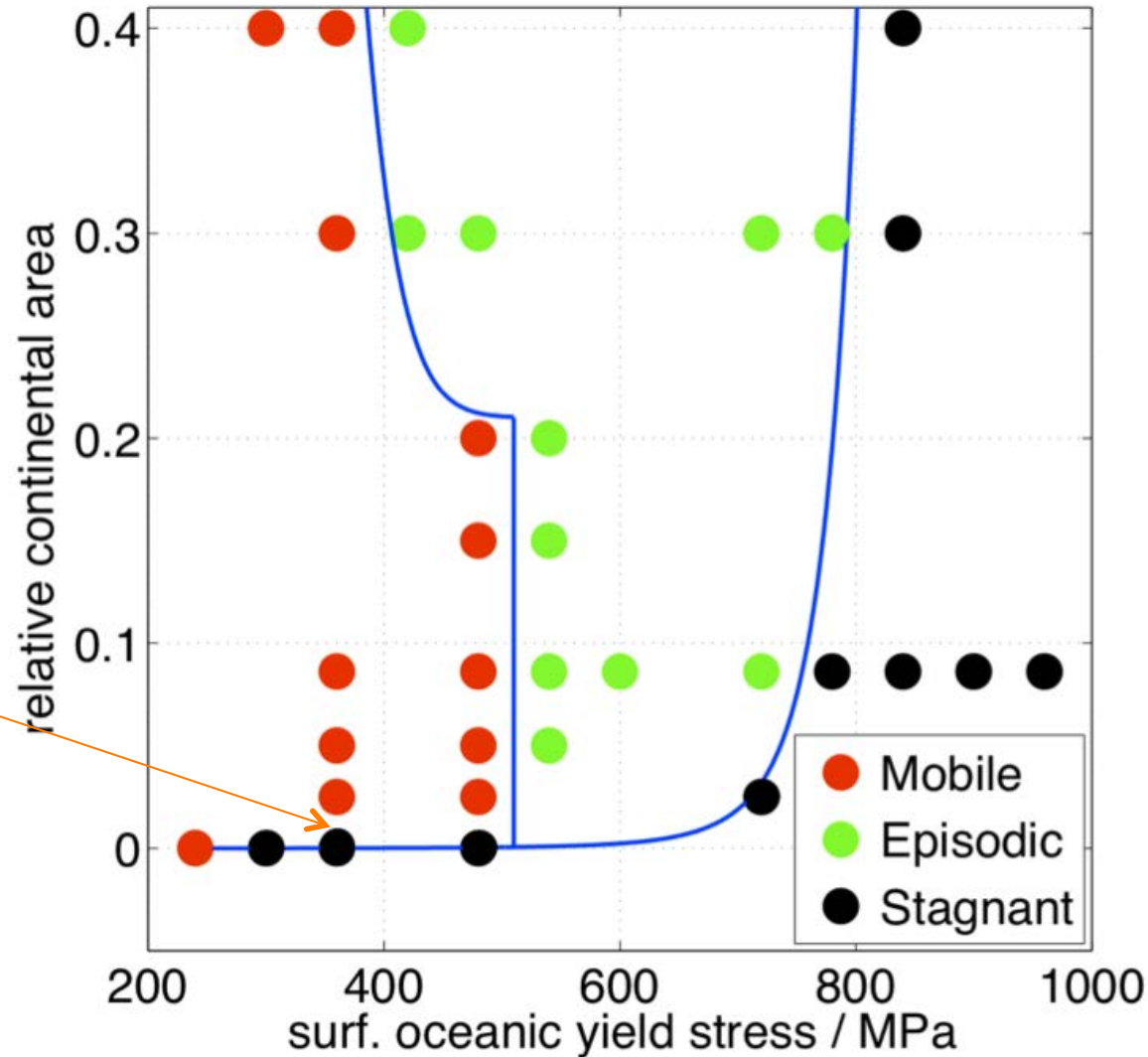
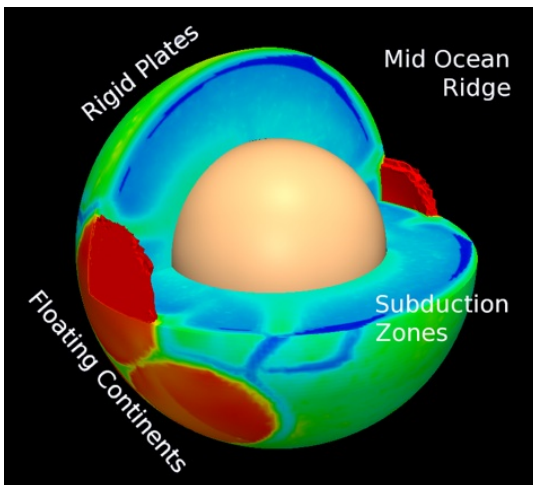
**far below laboratory values.**

**WHY?**



# Continents help plate tectonics!

Presence of continent allows plate tectonics at higher yield stress



Rolf and Tackley, GRL 2011



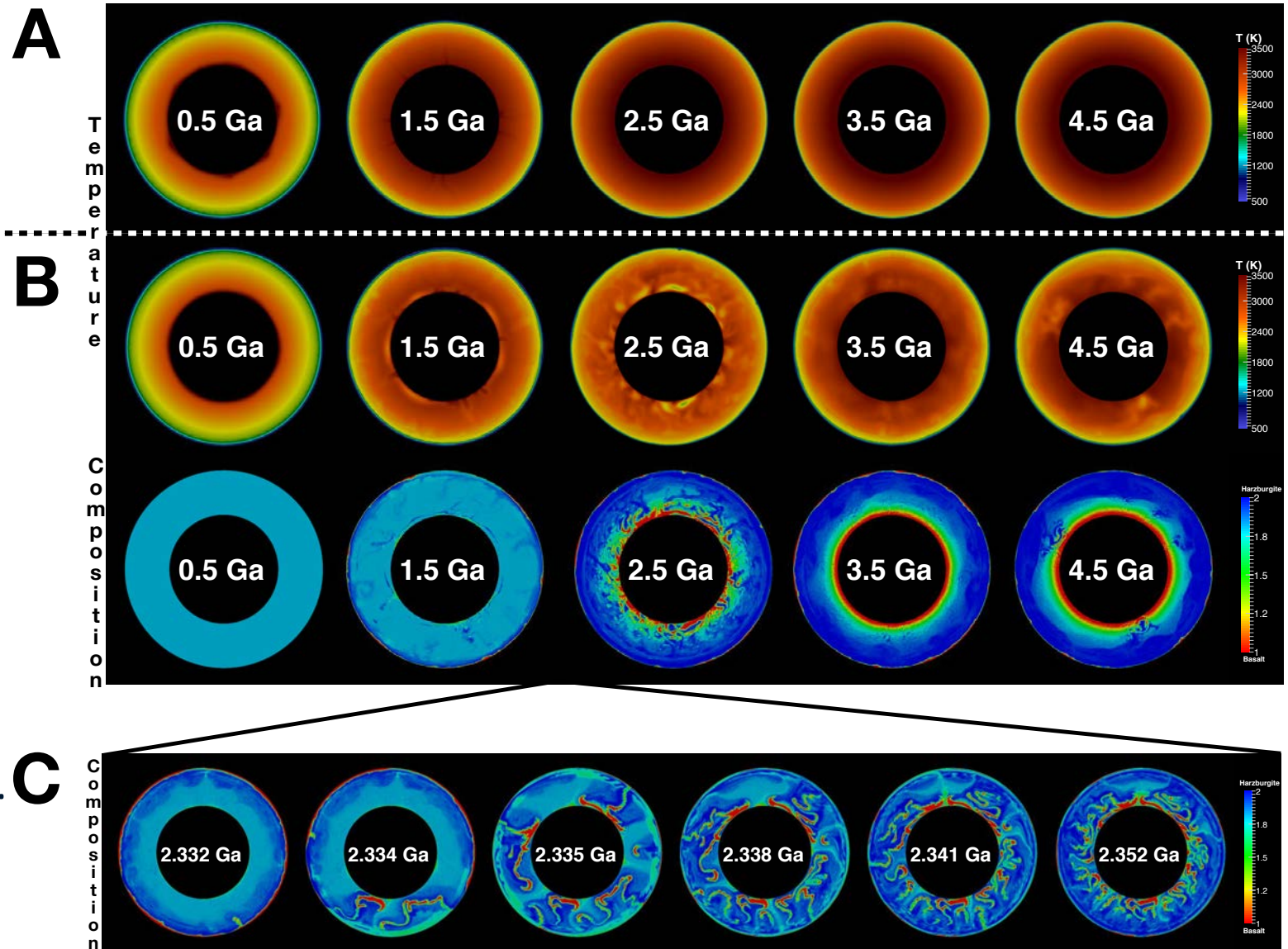
# Magmatism->crust helps plate tectonics

Purely thermal  
-> Stagnant

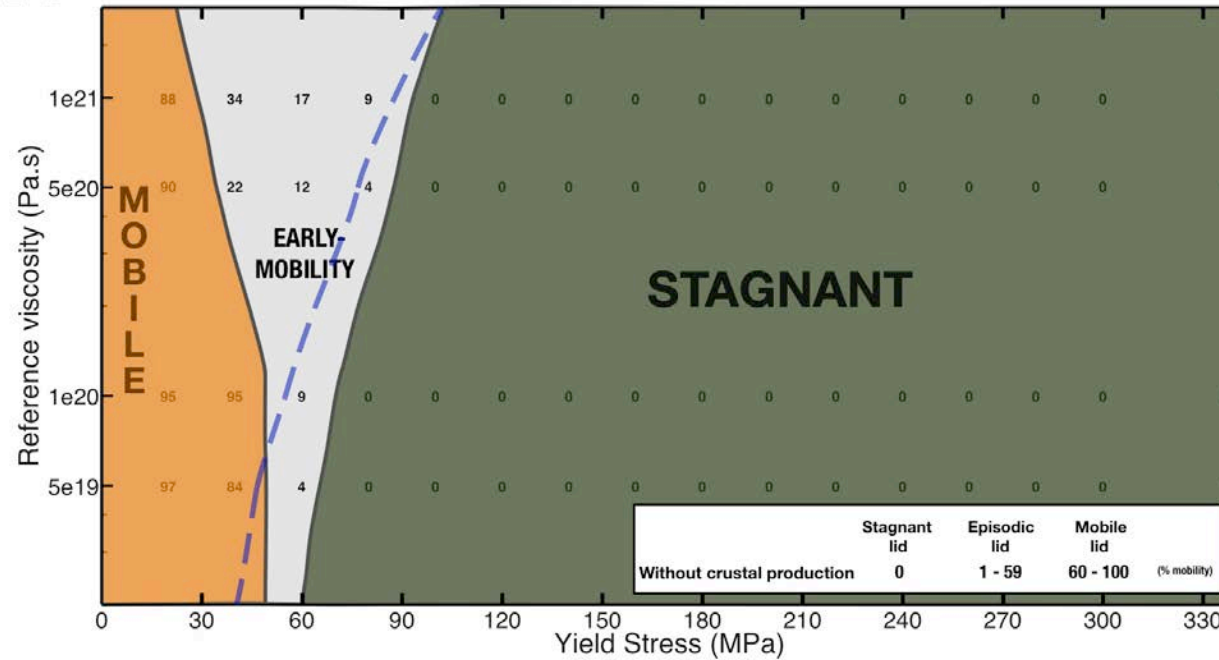
With magma &  
crust  
Episodic plate  
tectonics



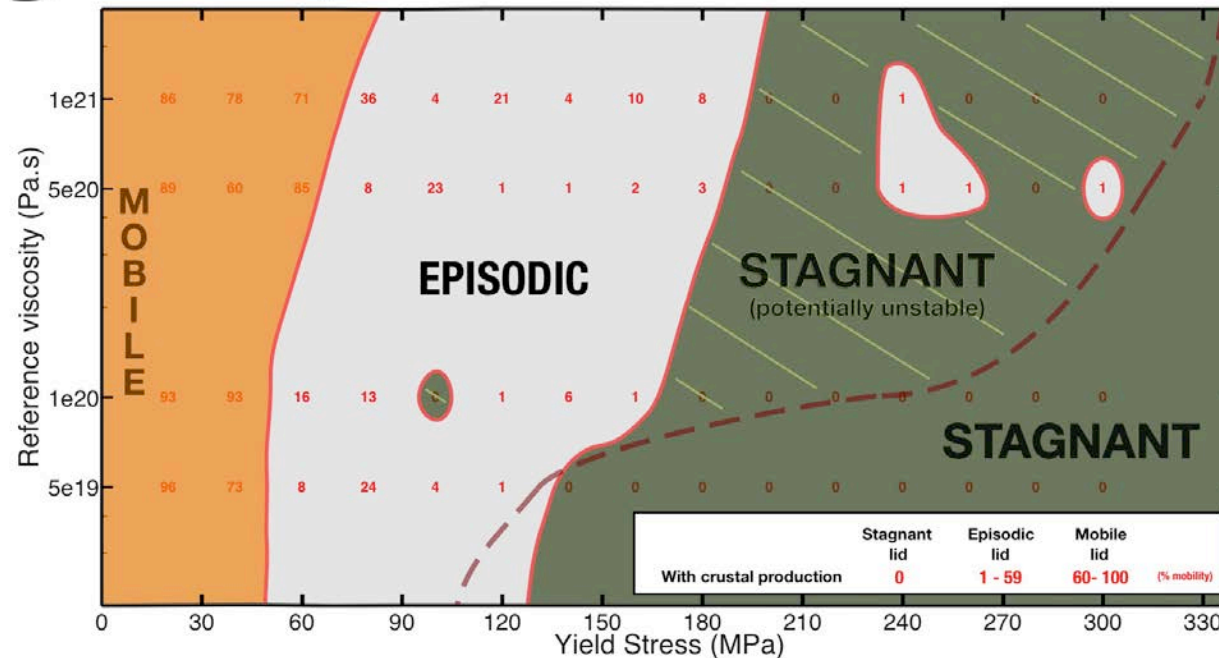
Diogo Lourenco A.  
Rozel & Tackley,  
EPSL 2016



## A - Without Melting and Crustal Production



## B - With Melting and Crustal Production

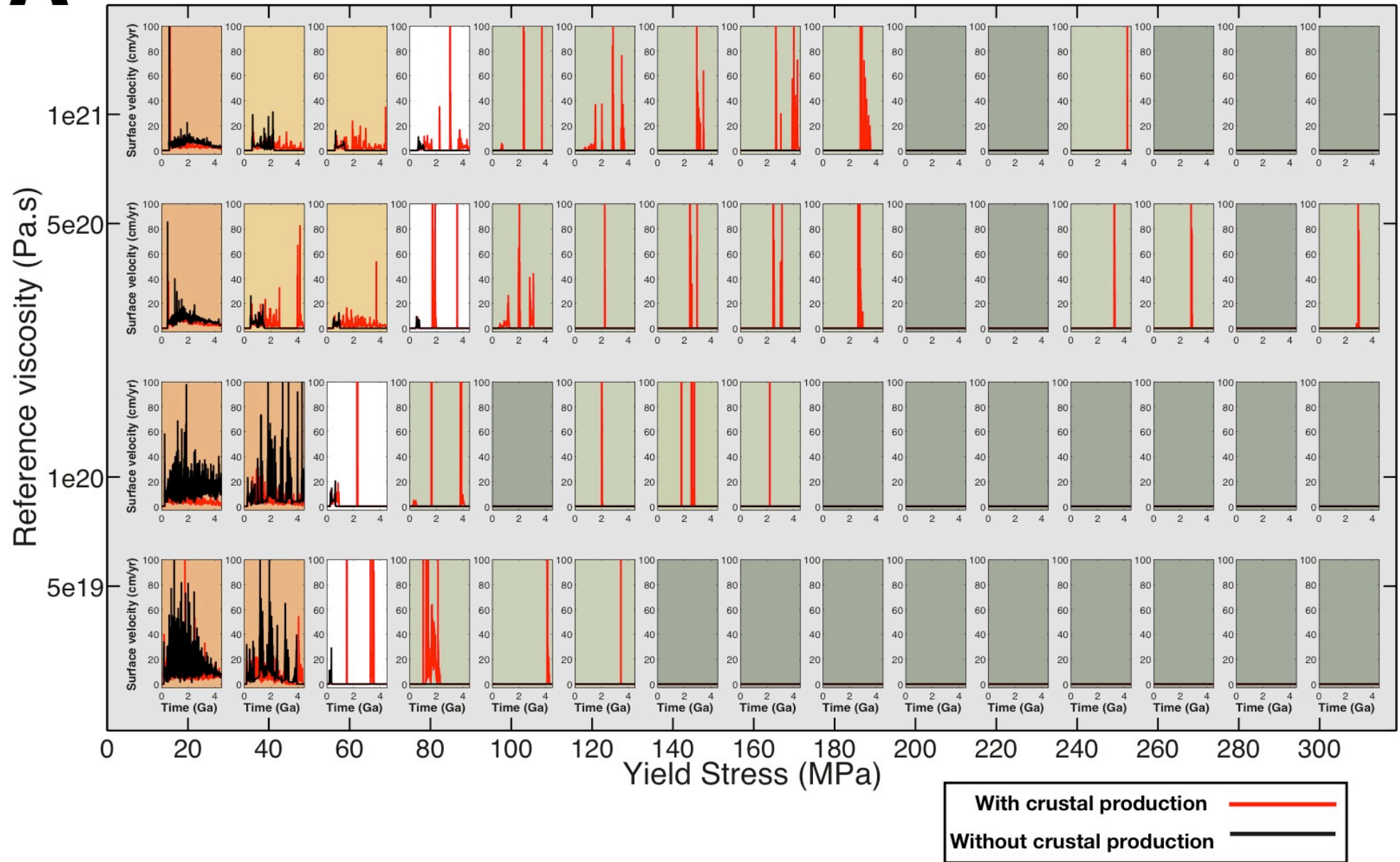


Lourenco et al.,  
EPSL 2016



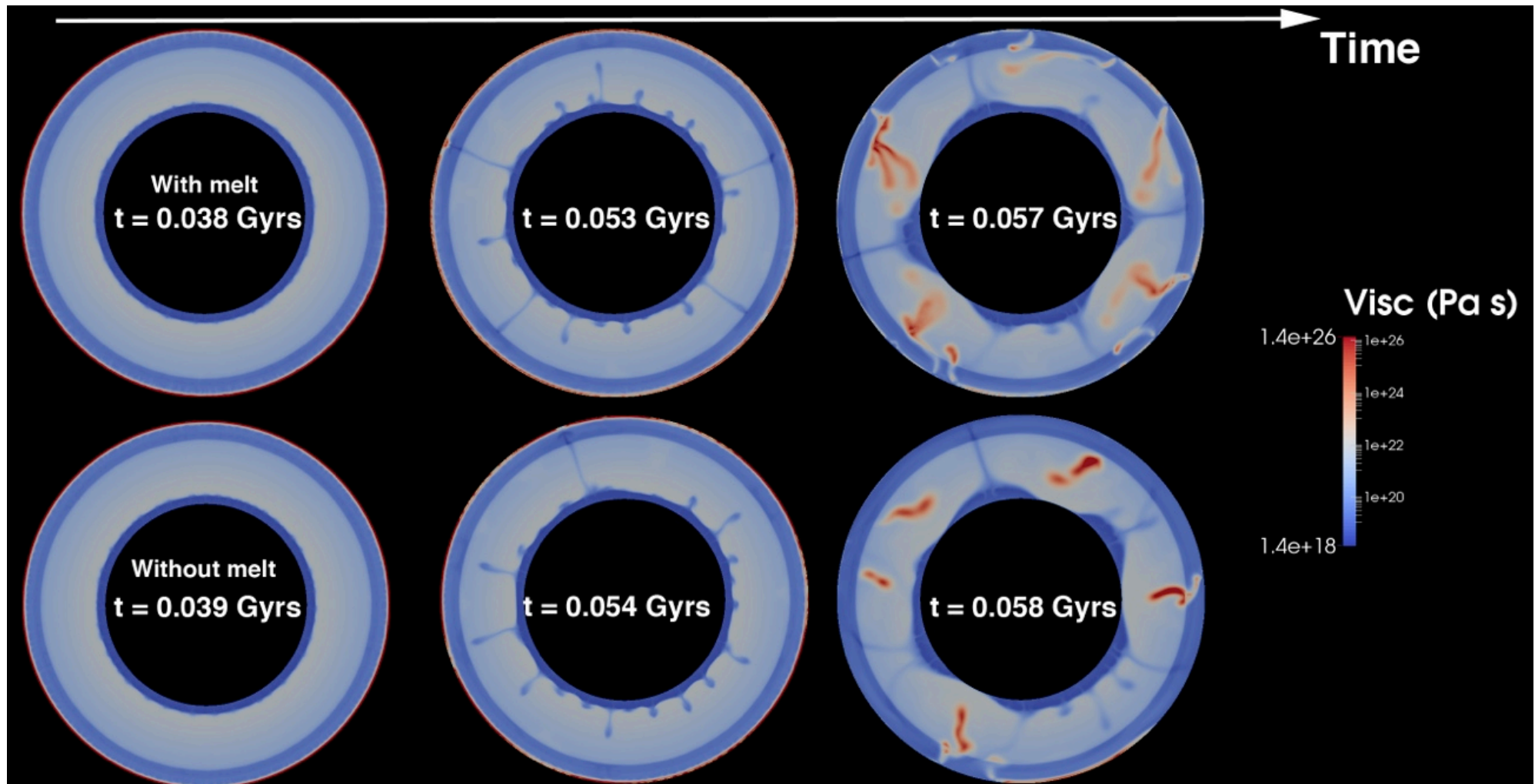
# Surface velocity vs. time

## A - Surface Velocity





# Initiation of first subduction by plumes

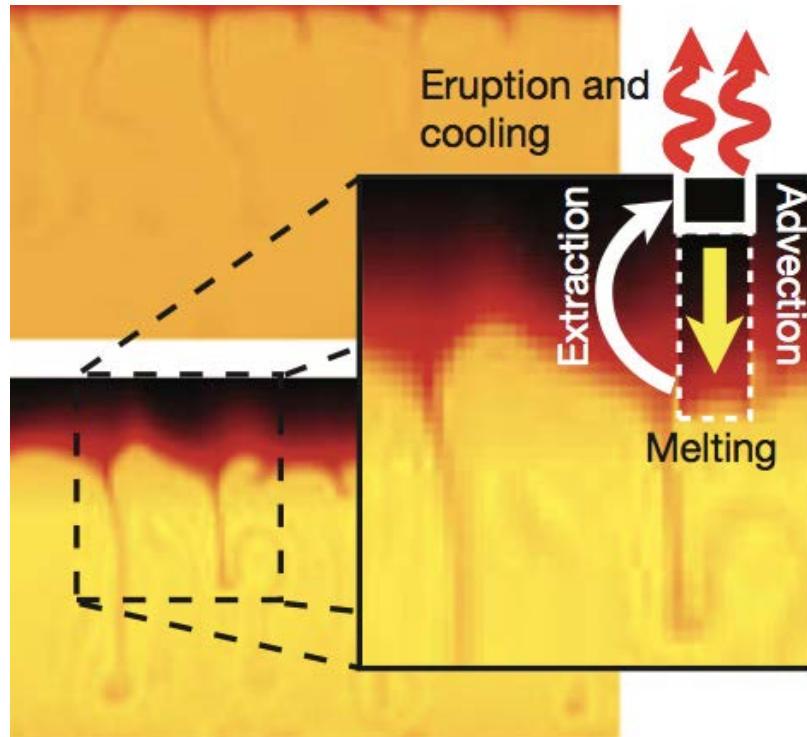


Nakagawa & Tackley, 2015

# Summary

- The production of laterally-varying crust assists plate tectonics on Earth-like planets
- Broad parameter range in which episodic behaviour occurs

# Extrusive heat pipe magmatism

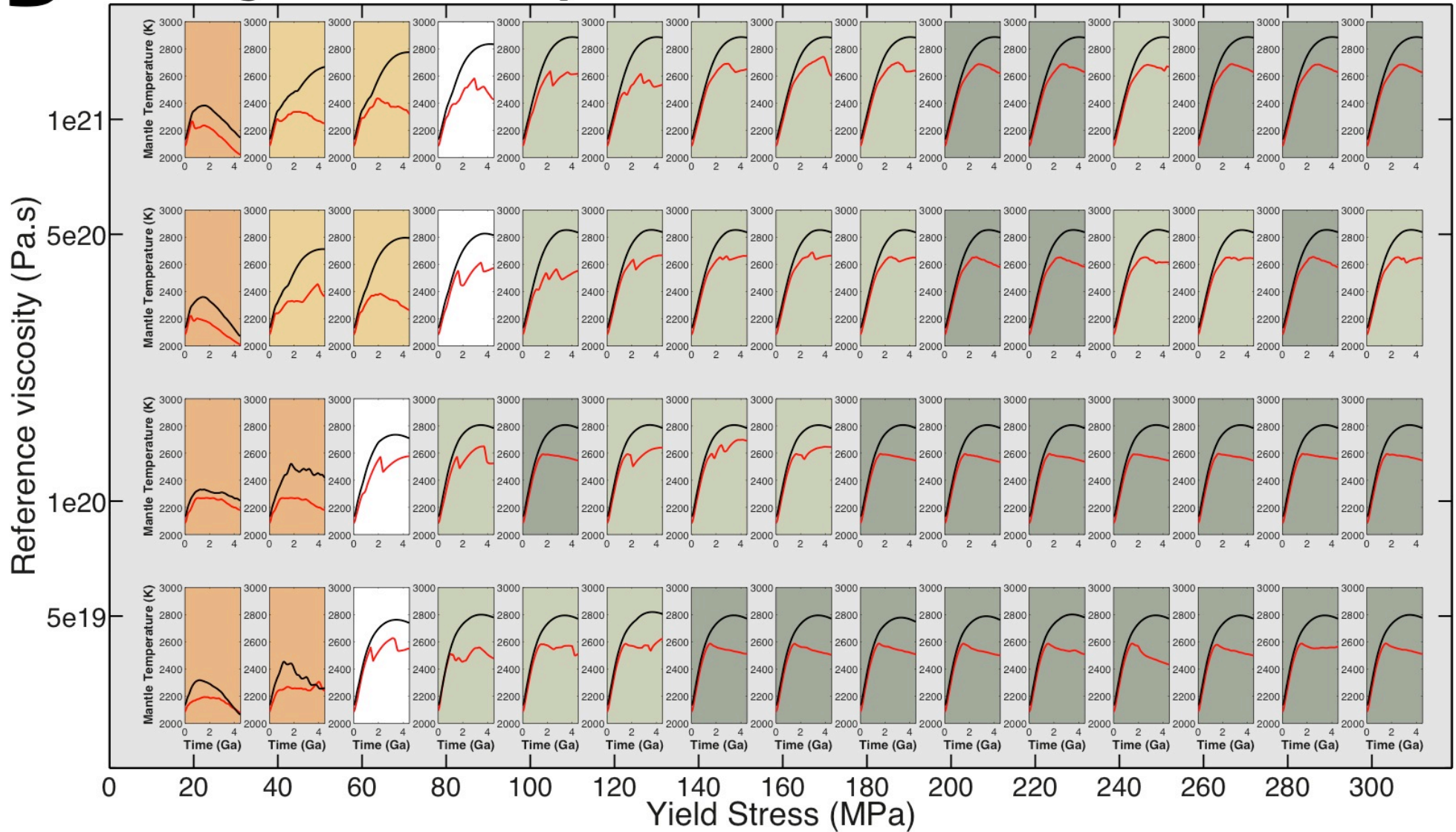
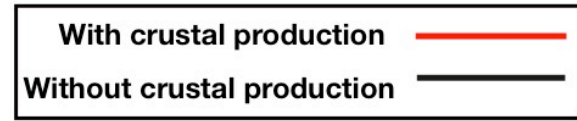


*(picture from Moore&Webb 2013)*

-> COLD, STRONG crust/lithosphere

# Temperature vs. time

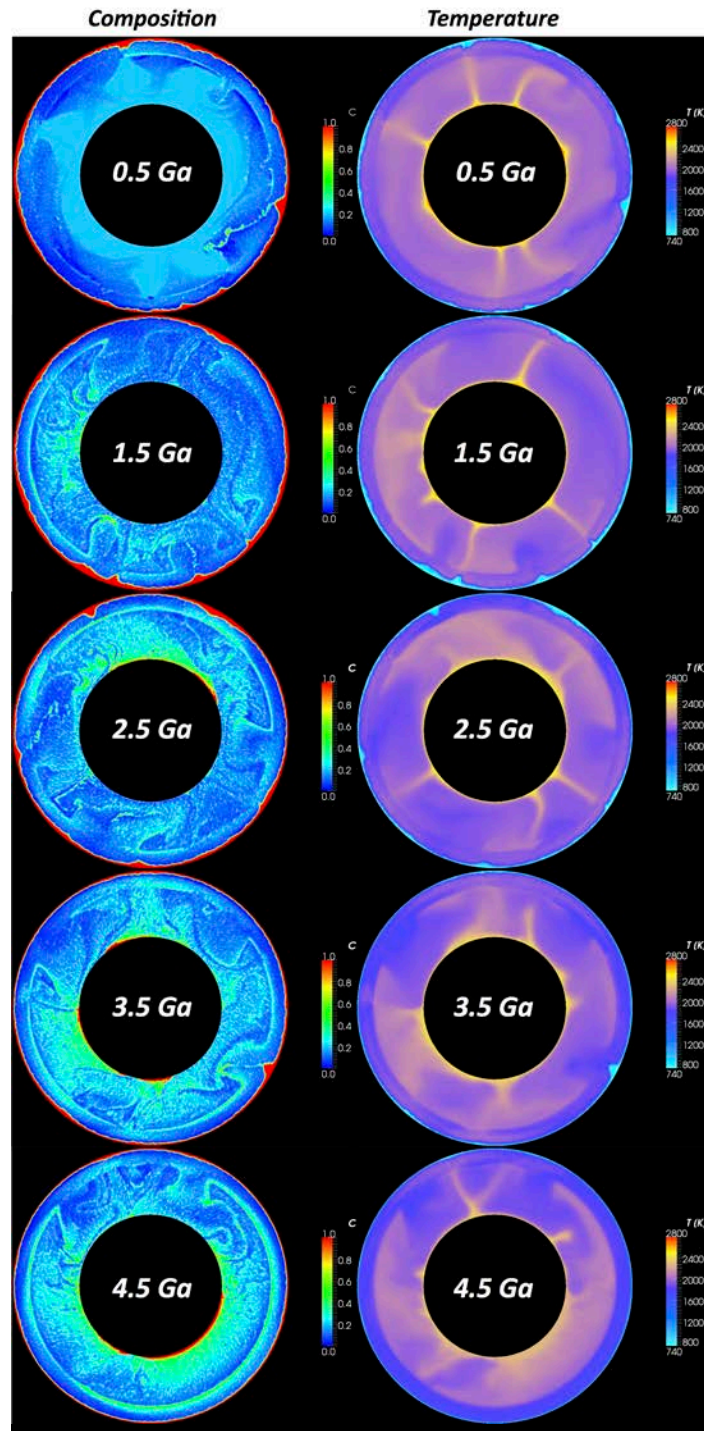
## B - Average Mantle Temperature



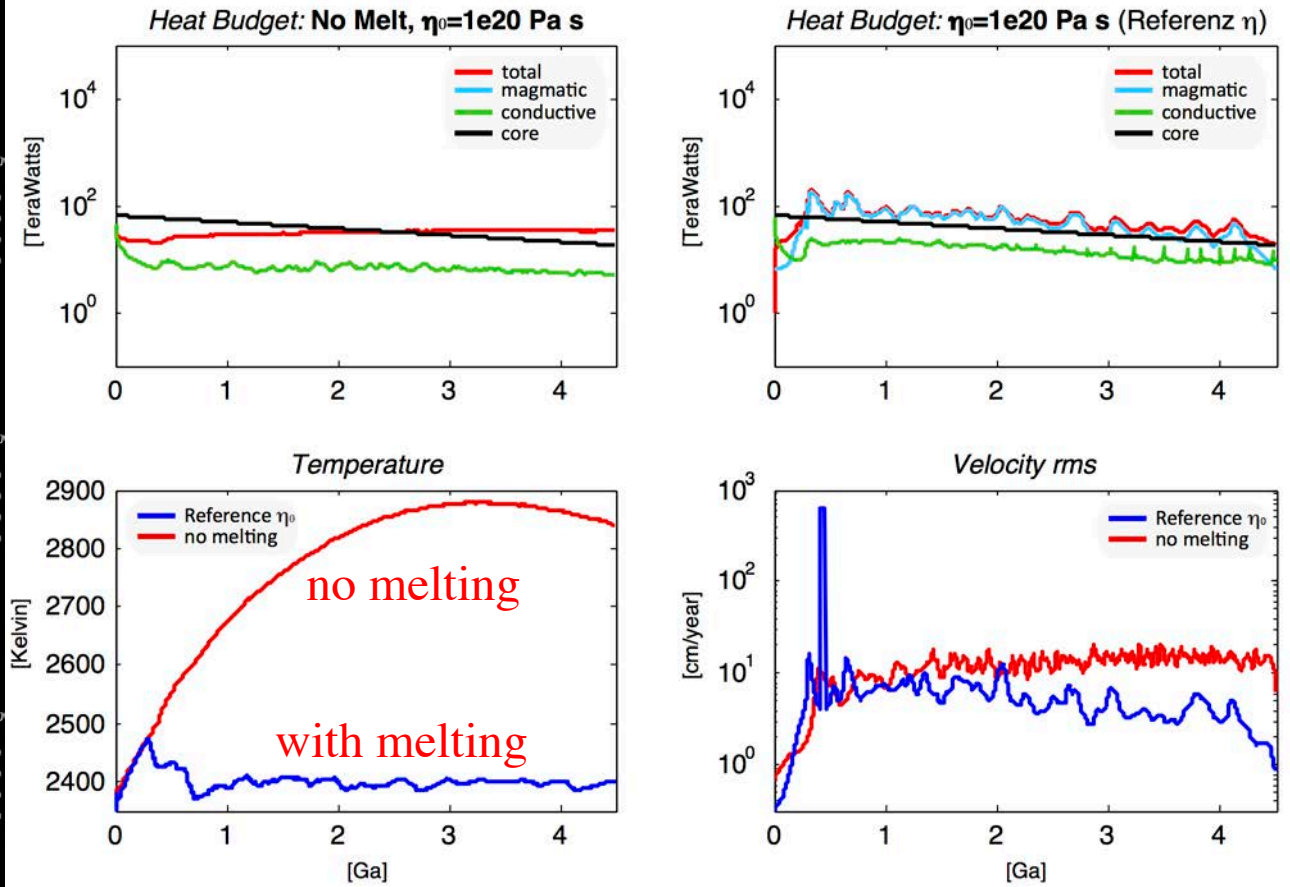


Reference Case:  $\eta=1e20$  Pa s

# Venus

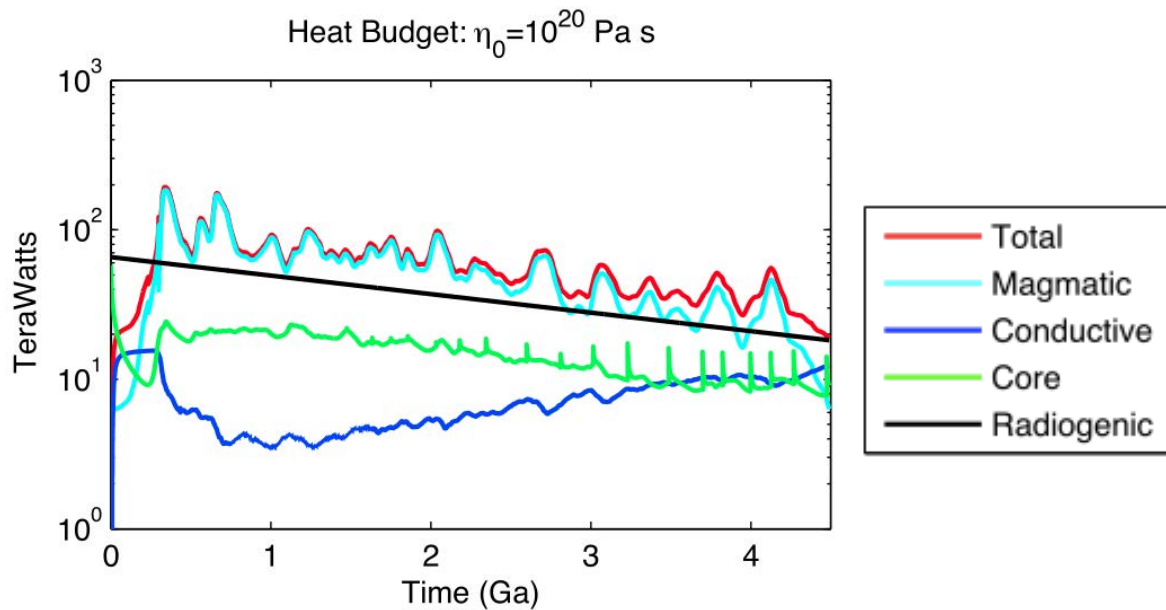


Thermal Evolution no melt vs melting case

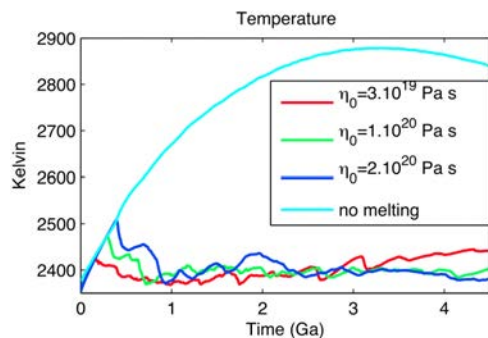


*Armann & Tackley, 2012 JGR*

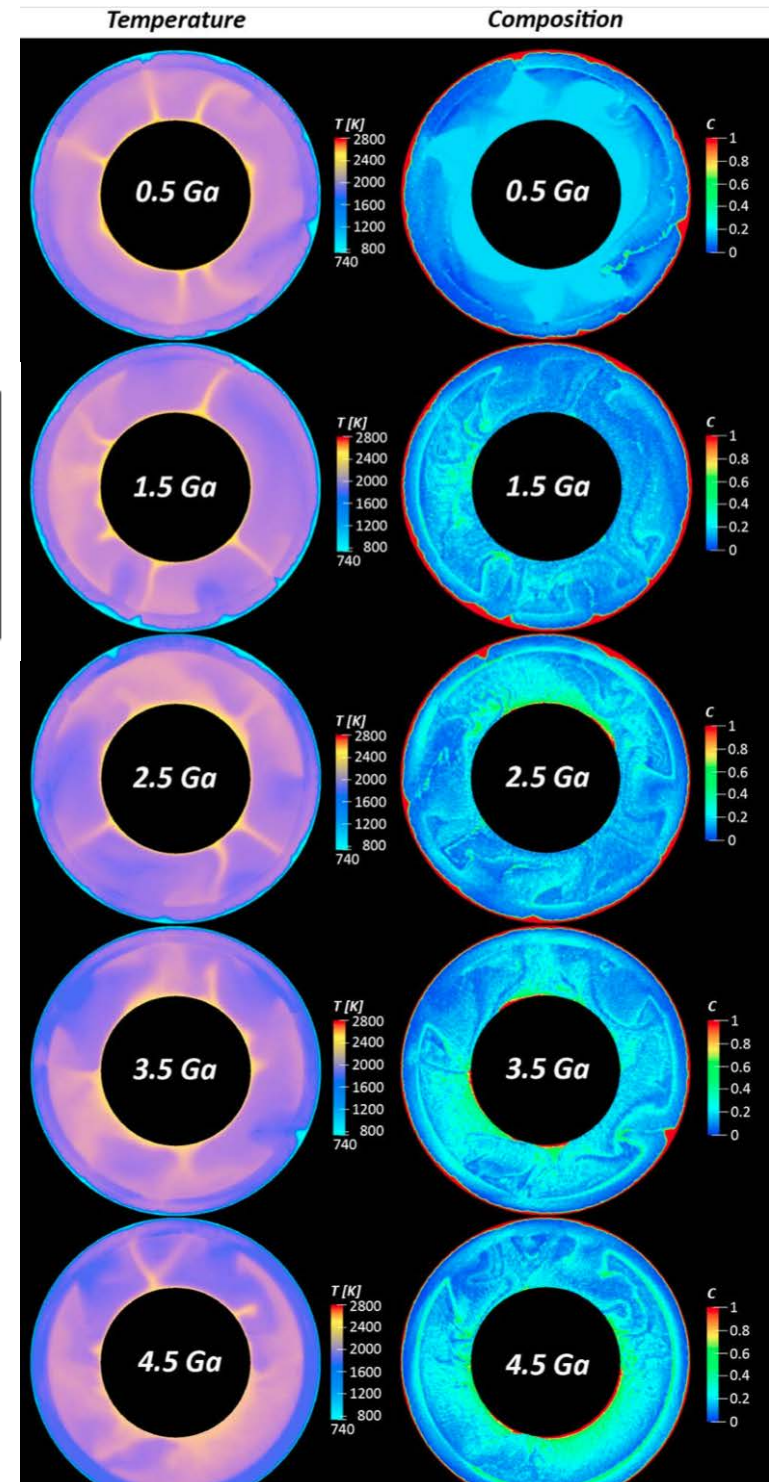
# Stagnant lid; extrusive



- ‘magmatic heat pipe’
- Cold crust/lithosphere
- High resurfacing rate (300 km/Gyr) does not match observations

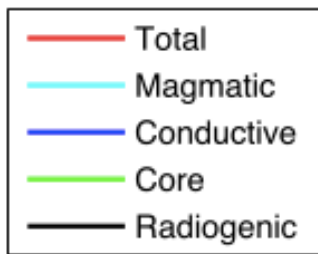
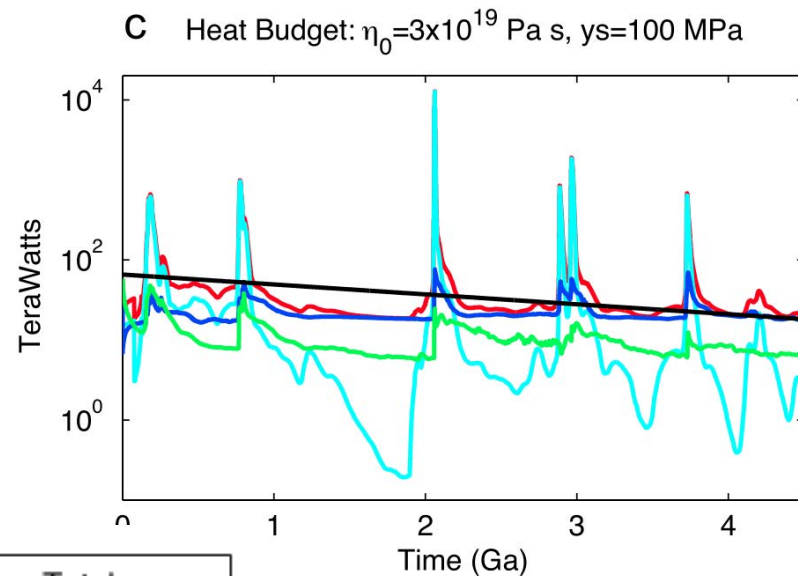


Magmatism => ‘thermostat’.  
No magmatism => high T,  
massive melting

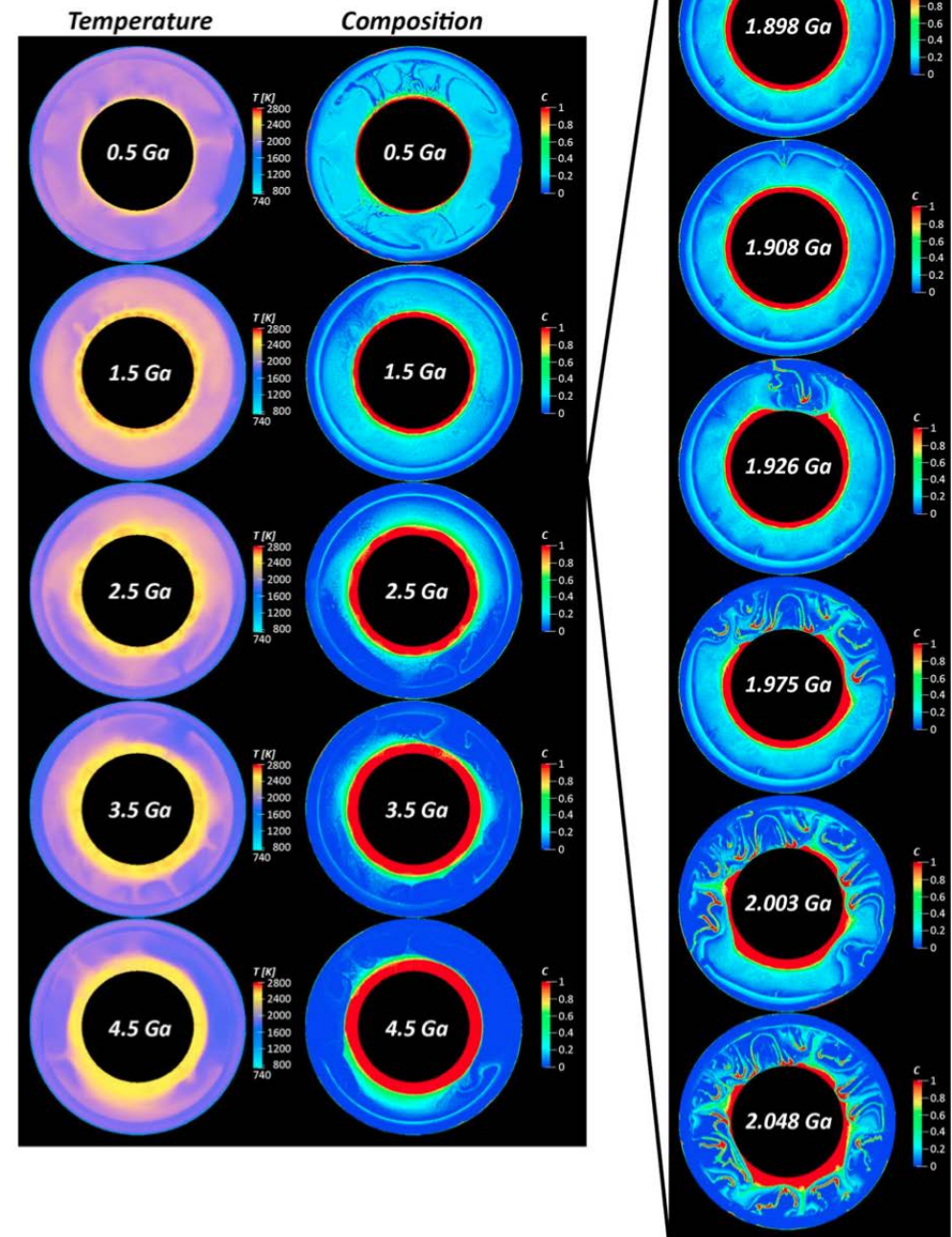




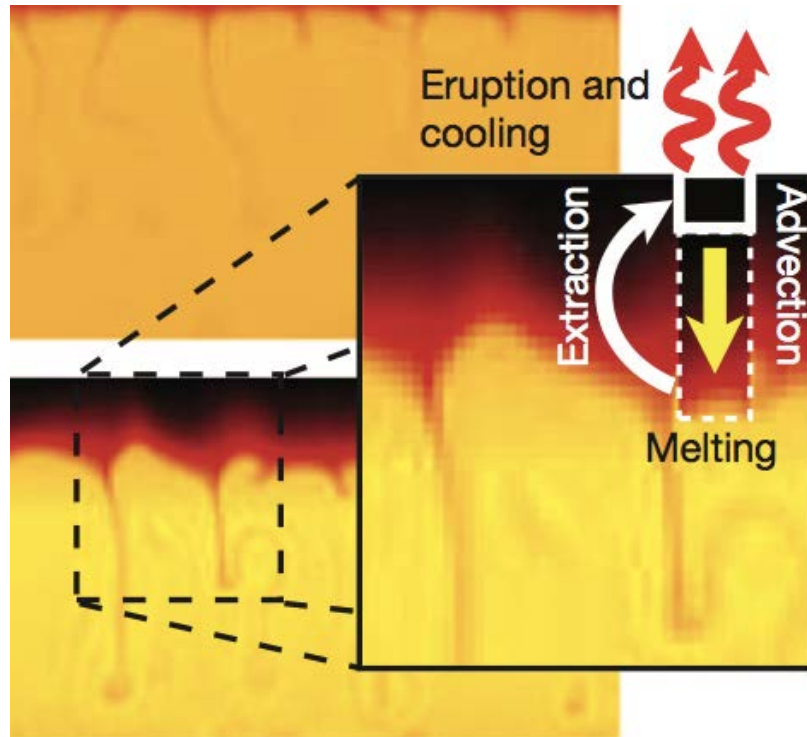
# Episodic lid; extrusive



- Episodic overturn due to plastic yielding
- Better matches surface observations



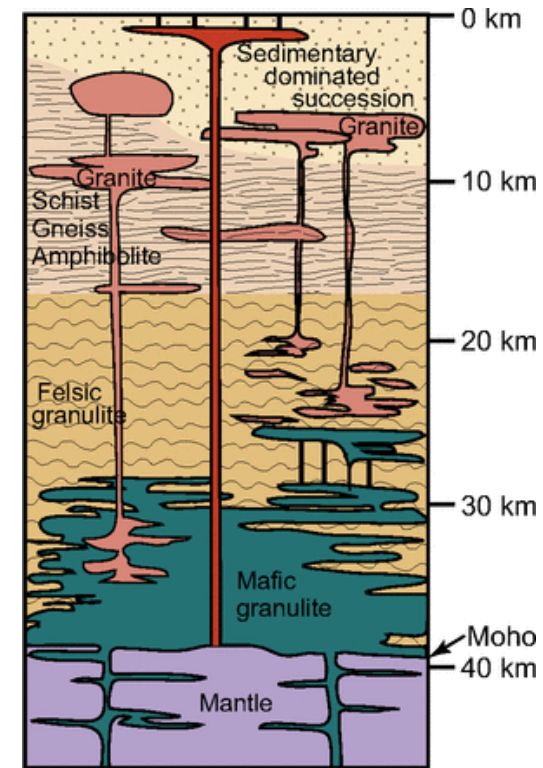
# Extrusive heat pipe magmatism



(picture from Moore&Webb 2013)

-> COLD, STRONG crust/lithosphere

But probably **most** magmatism is **intrusive**

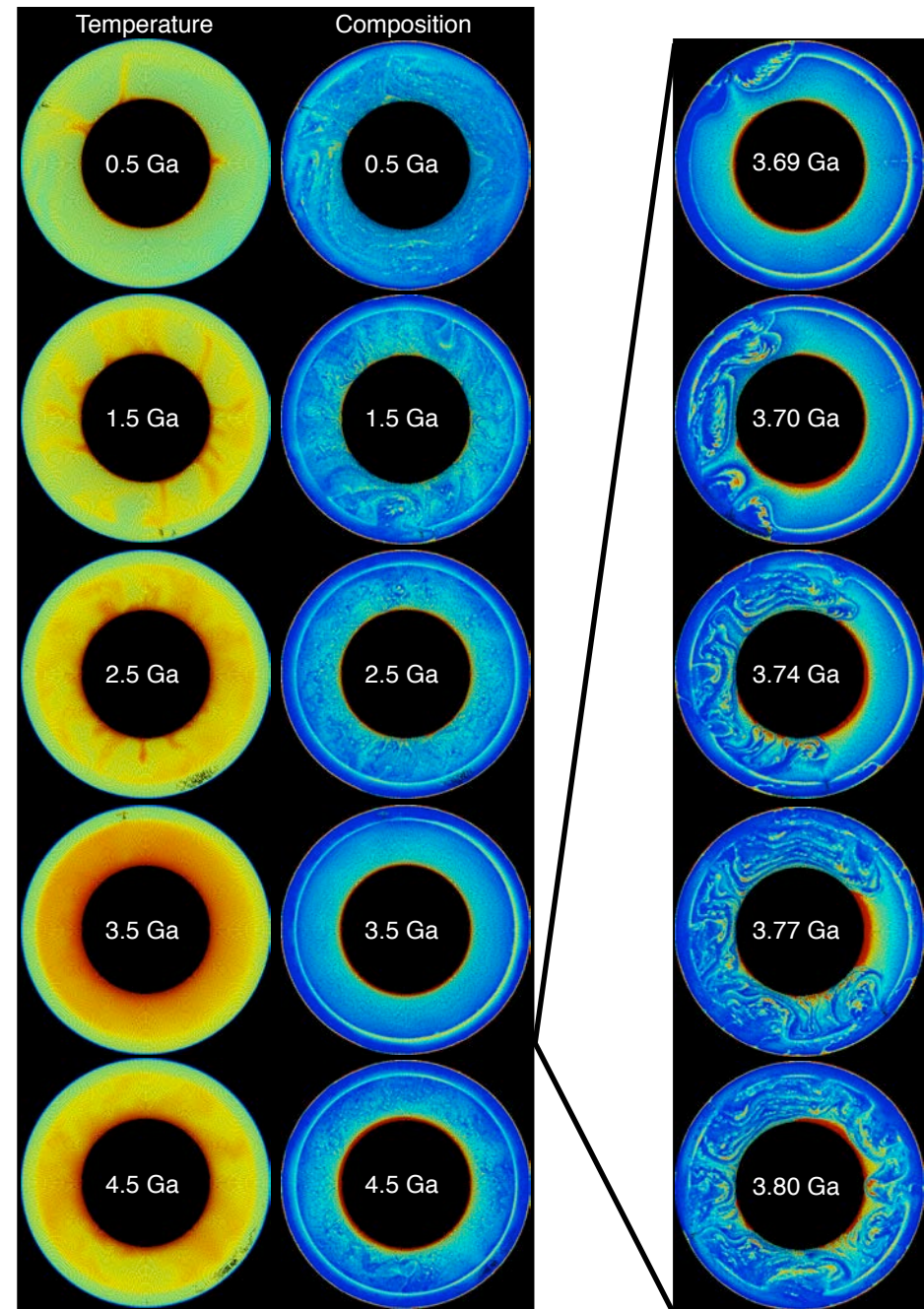
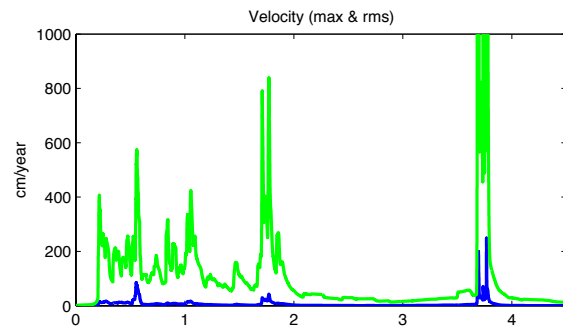
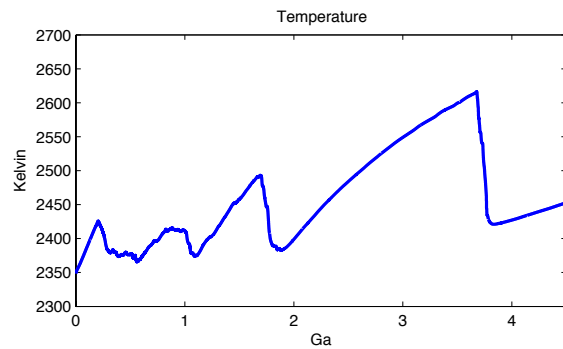
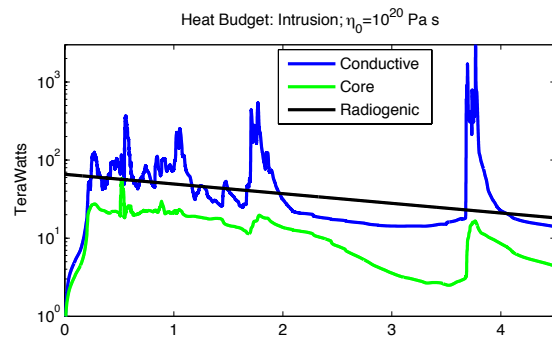


(picture from Cawood et al 2013)

-> WARM, WEAK crust/lithosphere



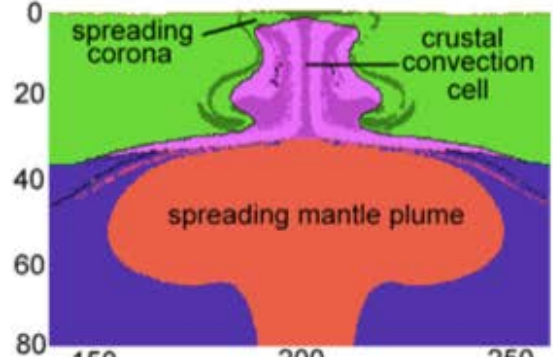
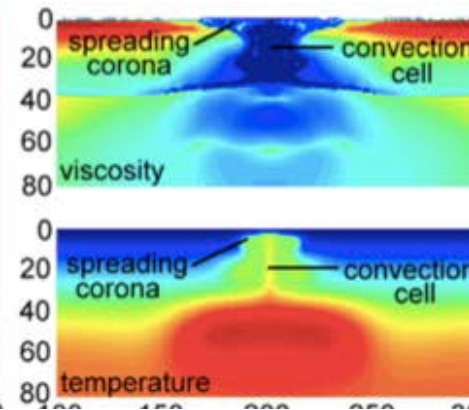
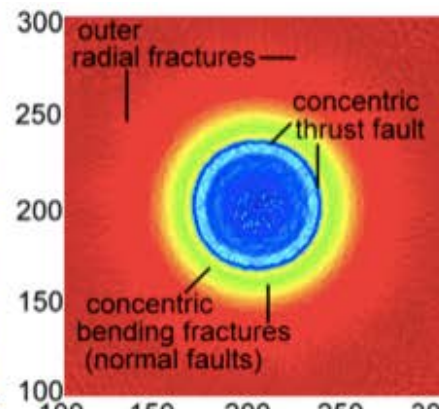
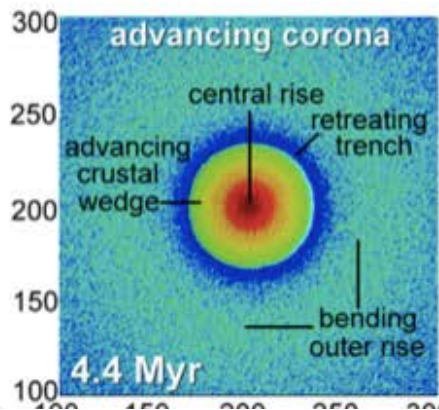
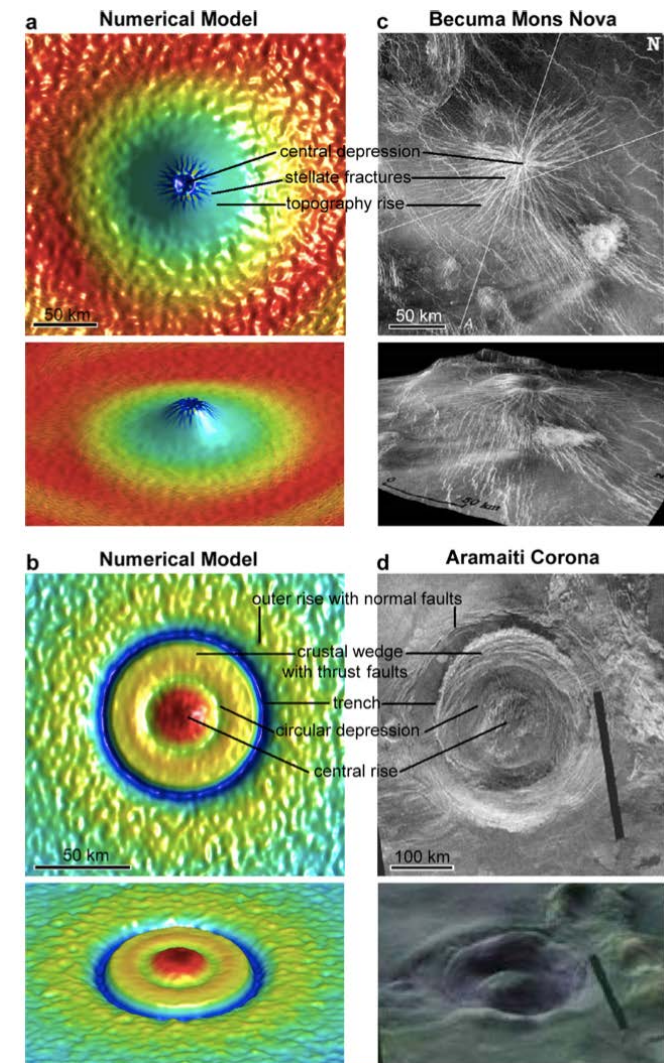
# Intrusive magmatism; no plastic yielding



# Intrusive magmatism; no plastic yielding

- Thin, weak crust & lithosphere
- Episodic overturn due to magmatic intrusion weakening lithosphere
- Also matches surface observations?

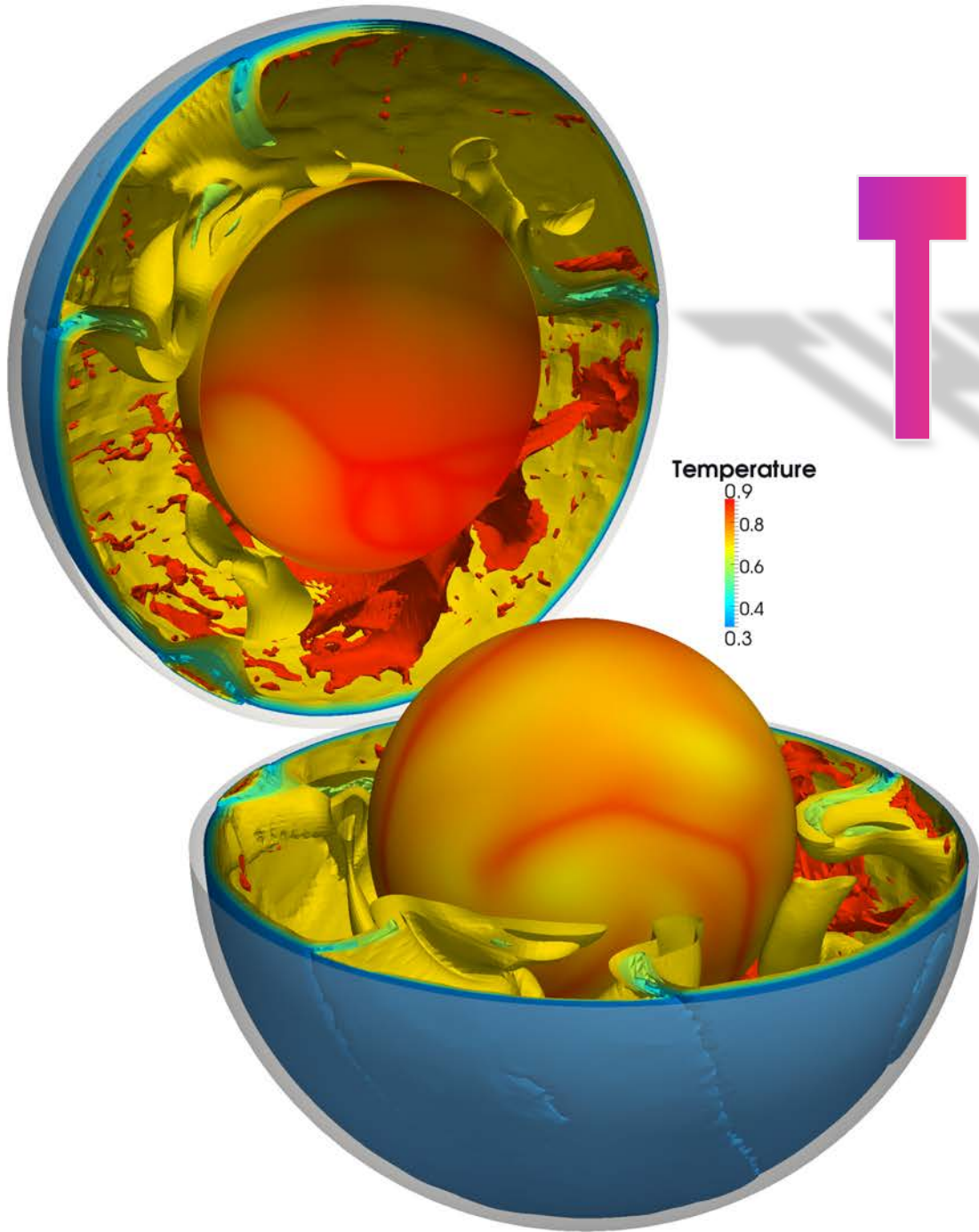
# Coronae & Novae caused by magmatic intrusion? – Gerya 2014 EPSL



# Conclusions

- Large-scale features of plate tectonics can be obtained with a simple description of plate boundaries
- Area-age distribution of oceanic lithosphere
- Size-number distribution of plates
- Melting and production of crust (oceanic or continental) play a big role in facilitating plate tectonics, as well as influencing other aspects of planetary evolution





THE END